$\mathcal{N}=8$ Self-Dual Supergravity

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Introduction Twistors and the Penrose Transform Twistors and Self-Dual Supergravity Conclusions and Outlook Supersymmetric Extensions

Twistor Space

- To begin with, consider complexified Minkowski 4-space $\mathbb{M} \cong \mathbb{C}^4$ coordinatised by $x^{\mu} \sim x^{\alpha \dot{\alpha}}$.
- A twistor is an object Z = (ω^α, π_ά) ∈ T satisfying the relation

$$\omega^{\alpha} = \mathbf{X}^{\alpha \dot{\alpha}} \pi_{\dot{\alpha}}.$$

Clearly, it's defined up to scalings. Thus twistors associated with \mathbb{M} are just points on $\mathbb{P} = \mathbb{P}^3 \setminus \mathbb{P}^1$ the latter being termed projective twistor space.

Twistor Space

- What do they describe geometrically?
- In fact, by virtue of $\omega^{\alpha} = \mathbf{x}^{\alpha \dot{\alpha}} \pi_{\dot{\alpha}}$ we find

$$\begin{array}{ccc} z = (\omega^{\alpha}, \pi_{\dot{\alpha}}) \in \mathbb{P} & \Longleftrightarrow & \text{isotropic 2-plane } \mathbb{C}_{2}^{2} \hookrightarrow \mathbb{M}: \\ & & x^{\alpha \dot{\alpha}} = x_{0}^{\alpha \dot{\alpha}} + \rho^{\alpha} \pi^{\dot{\alpha}} \\ \mathbb{P}_{x}^{1} \hookrightarrow \mathbb{P} & \iff & x = (x^{\alpha \dot{\alpha}}) \in \mathbb{M} \end{array}$$

Twistors and the Penrose Transform

- What's the meaning of these spheres $\mathbb{P}^1_x \hookrightarrow \mathbb{P}$?
- P¹_x corresponds to the projective null cone or the celestial sphere at x in real Minkowski space.

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Penrose Transform

- What is twistor space good for?
- Let's consider the following contour integral:

$$\phi(\mathbf{x}) = \frac{1}{2\pi i} \oint_{S^1} \mathsf{d} \, \pi^{\dot{\alpha}} \pi_{\dot{\alpha}} \, f(\mathbf{x}^{\beta \dot{\beta}} \pi_{\dot{\beta}}, \pi_{\dot{\beta}})$$

- Clearly, for this to be well-defined *f* must be homogeneous of degree -2 in $\pi_{\dot{\alpha}}$.
- It then follows that ϕ is a Klein-Gordon field

$$\Box \phi = \partial_{\alpha \dot{\alpha}} \partial^{\alpha \dot{\alpha}} \phi = \mathbf{0}.$$

• Being more precise, one has an isomorphism:

$$H^1(U', \mathcal{O}(-2)) \cong \{ \ker \Box \text{ on } U \subset \mathbb{M} \}, \text{ where } U' \subset \mathbb{P}$$

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Penrose Transform

Generally speaking, one can show that

$$H^1(U', \mathcal{O}(2h-2)) \cong H^0(U, \mathcal{Z}_h),$$

where \mathcal{Z}_h denotes the sheaf of solutions to the helicity *h* zero rest mass (z.r.m.) field equations.

This is called the Penrose transform.

 Thus, any solution to z.r.m. field equations can be represented by certain holomorphic "functions" on twistor space which are free of differential constraints.

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Generalisation

- So far, we've discussed linear field equations. Can we use the above ideas to learn more about non-linear equations?
- Yes, but how?
 - Replace space-time as background manifold by a twistor space and
 - try to reinterpret the physical theory in question on that space

 \Leftrightarrow

such that:

free

analytic data on twistor space

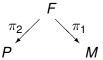
solutions

to the field equations on space-time

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Double Fibrations

 The central objects of our discussion are double fibrations of the form



where M, F and P are complex manifolds:

- *M* space-time
- F correspondence space
- P twistor space

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Twistor Correspondence: $P \stackrel{\pi_2}{\leftarrow} F \stackrel{\pi_1}{\rightarrow} M$

- Then we've a correspondence between *P* and *M*, i.e. between points in one space and subspaces of the other: $z \in P \iff \pi_1(\pi_2^{-1}(z)) \hookrightarrow M$ $\pi_2(\pi_1^{-1}(x)) \hookrightarrow P \iff x \in M$
- To jump ahead of our story a bit, what's the relation between *M* and *P* in later applications?
- Let x̂ := π₂π₁⁻¹(x) → P be compact for any x ∈ M. Assume further that H¹(x̂, N_{x̂|P}) = 0. Then M is taken to be the h⁰(x̂, N_{x̂|P})-dimensional family of deformations of x̂ inside P (which exists due to Kodaira's theorem).

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Twistor Transform: $P \stackrel{\pi_2}{\leftarrow} F \stackrel{\pi_1}{\rightarrow} M$

- Using the correspondence, we can transfer data given on *P* to data on *M* and vice versa.
- Take some analytic object Ob_P on P and transform it to an object Ob_M on M which will be constrained by some PDEs since $\pi_2^*Ob_P$ has to be constant along the fibers of $\pi_2 : F \to P$.
- Under suitable topological conditions, the maps

$$Ob_P \mapsto Ob_M$$
 and $Ob_M \mapsto Ob_P$

define a bijection between $[Ob_P]$ and $[Ob_M]$ (the objects in question will only be defined up to equivalence).

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Example A: Gauge Theory

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Penrose-Ward Transform

Consider the flat case again

Then there is a 1-1 correspondence between:

- holomorphic vector bundles *E*_ℙ on ℙ holomorphically trivial on any ℙ¹_x → ℙ,
- holomorphic vector bundles *E*_M on M equipped with a connection ∇ = d + A flat on each C²_z → M for z ∈ P, which implies ∇² = *F* = *F*⁺, i.e. self-dual YM fields.

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Example B: Gravity

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Penrose's Non-Linear Graviton

Consider a complex spin 4-manifold (M, [g]) equipped with a conformal structure $[g] = \{g' \sim g \mid g' = e^{\phi} g\}$. Hence $TM \cong S_+ \otimes S_-$



Then there is a 1-1 correspondence between complex 3-manifolds *P* and self-dual vacuum metrics on *M* such that:

- There is a holomorphic fibration $\pi: \mathbf{P} \to \mathbb{P}^1$.
- P has a 4-parameter family of sections each with normal bundle O(1) ⊕ O(1).
- There is a symplectic structure on the fibres of π with values in π*O(2).

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Supertwistors

- To discuss supersymmetry, we need supermanifolds.
- A supermanifold is a pair $(|M|, \mathcal{O}_M)$ such that:
 - |*M*| is a topological space and
 - \mathcal{O}_M is the structure sheaf describing the "superfunctions" on M with $\mathcal{O}_M \cong \Lambda^{\bullet} \mathcal{E}$ (locally).
- For the time being, simply think of them as manifolds coordinatised by a set of Z₂-graded coordinates (x^a, ηⁱ) and superfunctions are of the form

$$f(x,\eta) = f_0(x) + f_i(x)\eta^i + \cdots + f_{1\cdots m}(x)\eta^1 \cdots \eta^m.$$

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Example

• Consider $\mathbb{C}^{m+1|n} = (\mathbb{C}^{m+1}, \Lambda^{\bullet}\mathbb{C}^n)$ coordinatised by $(Z') = (z^a, \eta^i)$ and introduce the equivalence relation

$$Z' \sim t Z'$$
 for $t \in \mathbb{C}^*$.

This gives the projective superspace

$$\mathbb{P}^{m|n} = (\mathbb{P}^m, \Lambda^{\bullet}(\mathbb{C}^n \otimes \mathcal{O}(-1)).$$

• On $\mathbb{C}^{m+1|n}$, we may put

 $\Omega_0 \ = \ \textit{D}(\textit{d}\textit{Z}') \ \in \ \textit{H}^0(\mathbb{C}^{m+1|n},\textit{Ber}\,\Omega^1\mathbb{C}^{m+1|n}),$

i.e. when $dZ^{I} \mapsto dZ^{J}T_{J}^{I}$ we have $\Omega_{0} \mapsto \Omega_{0}Ber(T)$.

• Thus, it scales under $Z' \mapsto t Z' \Rightarrow dZ' \mapsto t dZ'$ as

 $\Omega_0 \mapsto t^{m+1-n}\Omega_0.$

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For *m* + 1 = *n*, it's scale invariant and thus descending to a holomorphic volume form Ω on ℙ^{n|n+1}, i.e.

$$\mathsf{Ber}(\mathbb{P}^{n|n+1}) := \mathsf{Ber}\,\Omega^1\mathbb{P}^{n|n+1} \cong \mathcal{O}.$$

- This leads us to the notion of formal Calabi-Yau supermanifolds. They are complex supermanifolds fulfilling the following equivalent statements:
 - existence of a nowhere vanishing holomorphic volume form,
 - having a trivial holomorphic Berezinian bundle,
 - having vanishing first Chern class.

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Supersymmetric Twistor Correspondence (Flat Case)

Consider right-chiral complexified Minkowski superspace
 M ≃ C^{4|2N} coordinatised by (x^{Aα̇}) = (x^{αα̇}, η^{iα̇}). A supertwistor is an object Z = (ω^A, π_{α̇}) ∈ T satisfying the relation

$$\omega^{\mathbf{A}} = \mathbf{X}^{\mathbf{A}\dot{\alpha}}\pi_{\dot{\alpha}}.$$

• Thus, $\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}}$ and

$$\mathbb{F} = \mathbb{C}^{4|2\mathcal{N}} \times \mathbb{P}^{1}$$
$$\pi_{2} \qquad \qquad \pi_{1}$$
$$\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}} \qquad \mathbb{M} = \mathbb{C}^{4|2\mathcal{N}}$$

Supersymmetric Penrose-Ward Transform

- Then there is a 1-1 correspondence between:
 - holomorphic vector bundles *E*_ℙ on ℙ holomorphically trivial on any ℙ¹_x → ℙ for *x* ∈ 𝔄,
 - holomorphic vector bundles *E*_M on M equipped with a connection ∇ which is flat on each C^{2|N}_z → M for *z* ∈ P: In this case we obtain *N*-extended self-dual SYM theory.
- Holomorphic vector bundles can be described within the Čech or Dolbeaut approaches. In the latter approach, we have (0, 1)-connection $\nabla^{0,1} : E \to E \otimes \Omega^{0,1}$ and E is holomorphic iff

$$F^{0,2} = (\nabla^{0,1})^2 = \bar{\partial}A^{0,1} + A^{0,1} \wedge A^{0,1} = 0.$$

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Supersymmetric Penrose-Ward Transform

• For $E_{\mathbb{P}}$ on $\mathbb{P} = \mathbb{P}^{3|4} \setminus \mathbb{P}^{1|4}$, the field equation $F^{0,2} = 0$ follows from varying

$$\boldsymbol{\mathcal{S}}[\boldsymbol{\mathcal{A}}^{0,1}] \;=\; \int \boldsymbol{\Omega} \wedge \text{tr} \left(\boldsymbol{\mathcal{A}}^{0,1} \wedge \bar{\partial} \boldsymbol{\mathcal{A}}^{0,1} + \frac{2}{3} \, \boldsymbol{\mathcal{A}}^{0,1} \wedge \boldsymbol{\mathcal{A}}^{0,1} \wedge \boldsymbol{\mathcal{A}}^{0,1} \right).$$

In addition, we have

$$\mathbf{A}^{0,1} = \mathbf{a} + \eta^{i} \psi_{i} + \eta^{i} \eta^{j} \phi_{ij} + \eta^{i} \eta^{j} \eta^{k} \epsilon_{ijkl} \psi^{l} + \eta^{1} \eta^{2} \eta^{3} \eta^{4} \mathbf{b},$$

which leads upon inserting into $F^{0,2} = 0$ to the $\mathcal{N} = 4$ gauge multiplet $(A, \Psi_i, \Phi_{ij}, \Psi^j, B)$ and the e.o.m. of $\mathcal{N} = 4$ self-dual SYM theory in 4*d*.

• Recall that the above action can be shown to be equivalent to the Siegel action.

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Self-Dual Supergravity

So what about the gravity sector?

In the following, I consider:

- Penrose's non-linear graviton construction from Ward's point of view,
- the supersymmetry extension thereof and
- the off-shell extension in the $\mathcal{N} = 8$ case.

Supersymmetric Non-Linear Gravitons Off-Shell $\mathcal{N}=8$ Self-Dual Supergravity

Ward's Theorem

There is a 1-1 correspondence between:

- complex 4-manifolds *M* with holomorphic metric *g* such that the trace-free Ricci tensor and anti-self-dual Weyl curvature vanish, but the scalar curvature is non-vanishing and
- complex 3-manifolds *P* with non-degenerate holomorphic contact structure containing a P¹ with normal bundle *O*(1) ⊕ *O*(1).

A non-degenerate holomorphic contact structure is a maximally non-integrable rank-2 subbundle *D* in $T^{1,0}P$. Think of *D* being defined as $D = \ker \tau$ with $\tau \in \Omega^{1,0}P$.

non-degeneracy $\iff \tau \wedge d\tau \neq 0.$

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How can one extend this to supergravity?

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Recap about Conformal Structures

Let M be a complex spin 4-manifold.

- A holomorphic conformal structure on *M* is an equivalence class [g] of metrics defined by [g] = {g' ~ g | g' = e^φ g}.
- Putting it differently, it's a line subbundle *L* in $\Omega^1 M \odot \Omega^1 M$.
- Since *M* is spin, we have $TM \cong S_+ \otimes S_-$. This isomorphism canonically yields a line subbundle $\Lambda^2 S^*_+ \otimes \Lambda^2 S^*_-$ in $\Omega^1 M \odot \Omega^1 M$.
- Actually, L and Λ²S^{*}₊ ⊗ Λ²S^{*}₋ may be identified, which thus yields an equivalent definition of a conformal structure.

Supersymmetric Extensions: The Setting

Let *M* be a right-chiral complex supermanifold of dimension $4|2\mathcal{N}$ coordinatised by $(x^{M\dot{\mu}}) = (x^{\mu\dot{\mu}}, \eta^{m\dot{\mu}})$.

- Suppose that *M* is split, i.e. $M = (|M|, \Lambda^{\bullet} \mathcal{E})$.
- Suppose also that *TM* ≅ *H* ⊗ *S*, where *H* is of rank 2|*N* and *S* of rank 2|0, respectively.
- As usual, let's introduce viel-beine

$$E^{A\dot{lpha}} = \mathrm{d} x^{M\dot{\mu}} E_{M\dot{\mu}}{}^{A\dot{lpha}}, \qquad E_{A\dot{lpha}} = E_{A\dot{lpha}}{}^{M\dot{\mu}} \partial_{M\dot{\mu}}.$$

• Let's also introduce a torsion-free connection ∇ on *TM* by

$$\nabla V^{A\dot{\alpha}} = \mathsf{d} V^{A\dot{\alpha}} + V^{B\dot{\alpha}} \omega_{B}{}^{A} + V^{A\dot{\beta}} \omega_{\dot{\beta}}{}^{\dot{\alpha}}$$

which preserves $TM \cong H \otimes S$.

Supersymmetric Extensions: The Setting

• Due to $TM \cong H \otimes S$, the curvature R of ∇ decomposes as

$$R_{A\dot{\alpha}}{}^{B\dot{\beta}} = \delta_{\dot{\alpha}}{}^{\dot{\beta}}R_{A}{}^{B} + \delta_{A}{}^{B}R_{\dot{\alpha}}{}^{\dot{\beta}}$$

The Ricci identity is then

$$\begin{split} [\nabla_{A\dot{\alpha}}, \nabla_{B\dot{\beta}}] V^{D\dot{\delta}} &= (-)^{p_{C}(p_{A}+p_{B})} V^{C\dot{\delta}} R_{A\dot{\alpha}B\dot{\beta}C}{}^{D} + \\ &+ (-)^{p_{D}(p_{A}+p_{B})} V^{D\dot{\gamma}} R_{A\dot{\alpha}B\dot{\beta}\dot{\gamma}}^{\dot{\delta}}. \end{split}$$

• The self-dual supergravity equations are given by

$$\begin{split} R_{A\dot{\alpha}B\dot{\beta}C}{}^{D} &= \epsilon_{\dot{\alpha}\dot{\beta}} \left\{ C_{ABC}{}^{D} - 2(-)^{p_{C}(p_{A}+p_{B})} \Lambda_{C\{A}\delta_{B]}{}^{C} \right\}, \\ R_{A\dot{\alpha}B\dot{\beta}\dot{\gamma}}{}^{\dot{\delta}} &= 2\Lambda_{AB}\delta_{(\dot{\alpha}}{}^{\dot{\delta}}\epsilon_{\dot{\beta})\dot{\gamma}}. \end{split}$$

Supersymmetric Non-Linear Gravitons Off-Shell $\mathcal{N}=8$ Self-Dual Supergravity

A Theorem

There is a 1-1 correspondence between:

- holomorphic complex solutions to *N*-extended self-dual supergravity with non-degenerate cosmological constant Λ_{AB} and
- complex supermanifolds *P* of dimension 3|*N* with a non-degenerate (even) contact structure *τ* and an embedded P¹ with normal bundle *O*(1) ⊗ C^{2|*N*}.

[MW, arXiv:0705.1422] [L. J. Mason and MW, arXiv:0706.1941]

 $\begin{array}{l} \mbox{Supersymmetric Non-Linear Gravitons} \\ \mbox{Off-Shell \mathcal{N}} = 8 \mbox{ Self-Dual Supergravity} \end{array}$

Sketch of Proof

We want to establish $P \stackrel{\pi_2}{\leftarrow} F \stackrel{\pi_1}{\rightarrow} M$. To do that, consider $F = P(S^*)$ over M with fibre coordinates $\pi_{\dot{\alpha}}$. We define the twistor distribution to be the rank-2|N distribution D_F on F given by

$$D_{F} := \left\langle \pi^{\dot{\alpha}} \boldsymbol{E}_{\boldsymbol{A}\dot{\alpha}} + \pi^{\dot{\alpha}} \pi_{\dot{\gamma}} \omega_{\boldsymbol{A}\dot{\alpha}\dot{\beta}}{}^{\dot{\gamma}} \frac{\partial}{\partial \pi_{\dot{\beta}}} \right\rangle.$$

Then

$$[D_F, D_F] \subset D_F \quad \iff \quad R_{A(\dotlpha B \doteta \dot\gamma \dot\delta)} = 0$$

and if so, we get $\pi_2 : \mathbf{F} \to \mathbf{P}$ with dim $\mathbf{P} = 3|\mathcal{N}$.

Supersymmetric Non-Linear Gravitons Off-Shell $\mathcal{N}=8$ Self-Dual Supergravity

Sketch of Proof

Since *F* is a \mathbb{P}^1 -bundle over *M*, we find:

$$\begin{array}{cccc} \mathbb{P}^1_x \hookrightarrow P & \Longleftrightarrow & x \in M \\ z \in P & \Longleftrightarrow & \text{isotropic } 2 | \mathcal{N} \text{-supermanifold in } M. \end{array}$$

The normal bundle of each \mathbb{P}^1_x is $\mathcal{O}(1) \otimes \mathbb{C}^{2|\mathcal{N}}$. The inverse construction uses a supersymmetric extension of Kodaira's deformation theory. Let's skip it ...

... instead, let's move on towards the self-dual Einstein condition.

 $\begin{array}{l} \mbox{Supersymmetric Non-Linear Gravitons} \\ \mbox{Off-Shell \mathcal{N}} = 8 \mbox{ Self-Dual Supergravity} \end{array}$

Sketch of Proof

So far, we have established $P \stackrel{\pi_2}{\leftarrow} F \stackrel{\pi_1}{\rightarrow} M$. Let's introduce a 1-form $\tilde{\tau}$ on F by

$$\widetilde{\tau} = \pi^{\dot{\alpha}} \nabla \pi_{\dot{\alpha}} = \pi^{\dot{\alpha}} \mathsf{d} \pi_{\dot{\alpha}} - \omega_{\dot{\alpha}}{}^{\dot{\beta}} \pi^{\dot{\alpha}} \pi_{\dot{\beta}}.$$

Clearly, $\widetilde{\tau}$ is of homogeneity 2. It descents down to P, i.e. $\widetilde{\tau}=\pi_2^*\tau,$ iff

$$D_{F} \sqcup \widetilde{\tau} = 0 = D_{F} \sqcup d\widetilde{\tau}.$$

The first condition is always fulfilled while the second one only iff the self-dual supergravity equations are satisfied. Furthermore, non-degeneracy of the contact structure corresponds to non-degeneracy of Λ_{AB} .

Supersymmetric Non-Linear Gravitons Off-Shell $\mathcal{N}=$ 8 Self-Dual Supergravity

From Ward to Penrose

So in summary, non-degenerate holomorphic contact structures on supertwistor spaces correspond to solutions to the self-dual supergravity equations with non-zero cosmological constant.

What about the zero case?

In order for that to happen, one considers maximally degenerate contact structures on *P*. In particular, that means that $[D, D] \subset D$ for $D = \ker \tau$, i.e. *D* is integrable. Thus, we get a rank-2| \mathcal{N} foliation with base space \mathbb{P}^1 , i.e.

 $\pi: \boldsymbol{P} \to \mathbb{P}^1.$

Finite Deformations of $\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}}$

We want to describe a twistorial off-shell formulation of $\mathcal{N}=8$ self-dual supergravity.

But how?

In the following, I choose as background

- flat supertwistor space $\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}}$ with the standard $\bar{\partial}_0$ -operator and holomorphic contact structure τ_0 and
- consider finite complex and contact structure deformations on ℙ to get P with ∂ and τ.

This will give a way to obtain an off-shell formulation for $\mathcal{N} = 8$.

Supersymmetric Non-Linear Gravitons Off-Shell $\mathcal{N}=$ 8 Self-Dual Supergravity

Finite Deformations of $\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}}$

Consider $\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}}$ with coordinates $(z^{I}) = (z^{a}, \eta^{i}) = (\omega^{\alpha}, \pi_{\dot{\alpha}}, \eta^{i})$. Let's introduce a Poisson structure on homogeneous functions *f* and *g*

$$[f,g] := (-)^{p_I} \partial_I f \, \omega^{IJ} \, \partial_J g.$$

Let's also introduce a (0, 1)-form $h = d\bar{z}^{\bar{a}}h_{\bar{a}}$ of homogeneous degree 2 in z^{I} and 0 in $\bar{z}^{\bar{l}}$ holomorphic in η^{i} .

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Finite Deformations of $\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}}$

We then define the distribution $T^{0,1}P$ of anti-holomorphic tangent vectors on P by

$$T^{0,1}P := \langle \bar{D}_{\bar{l}} \rangle := \left\langle \frac{\partial}{\partial \bar{z}^{\bar{a}}} + (-)^{p_l} \frac{\partial h_{\bar{a}}}{\partial z^l} \omega^{IJ} \frac{\partial}{\partial z^J}, \frac{\partial}{\partial \bar{\eta}^{\bar{l}}} \right\rangle$$

This is to be understood as a finite perturbation of the standard complex structure on \mathbb{P} with $\bar{\partial}_0 = d\bar{z}^{\bar{l}}\bar{\partial}_{\bar{l}}$.

Since the (1,0)-forms are $Dz^{I} = dz^{I} + \omega^{IJ}\partial_{J}h$ and $\tau_{0} = dz^{I}z^{J}\omega_{JI}$, the deformed contact structure τ is

$$\tau = Dz^{I}z^{J}\omega_{JI} = dz^{I}z^{J}\omega_{JI} + z^{J}\underbrace{(-)^{p_{I}}\omega_{JI}\omega^{IK}}_{= \delta_{J}^{K}}\partial_{K}h = \tau_{0} + 2h.$$

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Finite Deformations of $\mathbb{P} = \mathbb{P}^{3|\mathcal{N}} \setminus \mathbb{P}^{1|\mathcal{N}}$

Then

$$\begin{split} [T^{0,1}P, T^{0,1}P] \subset T^{0,1}P & \iff & \omega^{IJ}\partial_J \left(\bar{\partial}_0 h + \frac{1}{2}[h,h] \right) = 0, \\ & \bar{\partial}\tau = 0 & \iff & \bar{\partial}_0 h + \frac{1}{2}[h,h] = 0, \end{split}$$

Thus, integrability of the complex structure follows from the holomorphy of the contact structure.

For $\mathcal{N} = 8$, $\overline{\partial}_0 h + \frac{1}{2}[h, h] = 0$ follows from

$$S[h] = \int \Omega \wedge \left(h \wedge \overline{\partial}_0 h + \frac{1}{3}h \wedge [h,h]\right),$$

since the homogeneity of *h* is 2, of $[\cdot, \cdot]$ is -2 and of Ω is -4

Action Principle for $\mathcal{N} = 8$ Self-Dual Supergravity

So we have

$$S[h] = \int \Omega \wedge \left(h \wedge \overline{\partial}_0 h + \frac{1}{3}h \wedge [h, h]\right).$$

Depending on the degeneracy of the Poisson structure ω , one can describe different self-dual supergravities:

- rank $\omega = 4 | \mathcal{N} \iff R$ -symmetry maximally gauged to $SO(\mathcal{N})$ & non-zero cosmological constant
- rank $\omega = 4|r \iff R$ -symmetry gauged to some $H \subset SO(\mathcal{N})$ & non-zero cosmological constant
- rank ω = 2|r ⇐⇒ R-symmetry gauged to some H ⊂ SO(N) & zero cosmological constant

Note that since a CS action depends on a chosen background, diffeomorphism invariance (on supertwistor space) is explicitly broken.

Supersymmetric Non-Linear Gravitons Off-Shell $\mathcal{N}=$ 8 Self-Dual Supergravity

Can one find a covariant formulation making diffeomorphism invariance manifest?

Covariant Approach for $\mathcal{N} = 0$: A Theorem

Suppose that on a (smooth) manifold *P* of dimension 4n + 2 we are given a complex line bundle $L^* \subset \mathbb{C}T^*P$, represented by a complex 1-form τ defined up to complex rescalings. Suppose further that

•
$$\tau \wedge (d\tau)^{n+1} = 0$$
 and $\tau \wedge (d\tau)^n \neq 0$ and

• ker{
$$\tau \wedge (d\tau)^n$$
} $\cap \overline{\ker\{\tau \wedge (d\tau)^n\}} = \{0\}.$

Then there is a unique integrable almost complex structure for which τ is proportional to a non-degenerate holomorphic contact structure.

[L. J. Mason and MW, arXiv:0706.1941]

Covariant Approach for $\mathcal{N} = 0$: An Action

In the twistor context, n = 1 and P is topologically $\mathbb{R}^4 \times S^2$ and $c_1(L) = 2$. Then, the field equation $\tau \wedge (d\tau)^2 = 0$ follows from

$$S[b, \tau] = \int b \wedge \tau \wedge (\mathsf{d}\tau)^2,$$

where $b \in \Omega^1 P \otimes (L^*)^3$ is a Lagrange multiplier.

On-shell, L becomes $\mathcal{O}(2)$ and $b \in H^1(P, \mathcal{O}(-6))$.

Via the Penrose transform *b* corresponds to a helicity -2 field propagating in the 4-dimensional self-dual background determined by τ .

Conclusions and Outlook

What we've got:

- We saw how Ward's extension of Penrose's non-linear graviton construction needs to be extended to self-dual supergravity.
- For N = 8, we found a Chern-Simons-like off-shell formulation of the theory on supertwistor space which was dependent on a choice of background.
- We made progress towards an invariant off-shell formulation for $\mathcal{N} = 0$.

[MW, arXiv:0705.1422] [L. J. Mason and MW, arXiv:0706.1941]

Conclusions and Outlook

What's next:

- Invariant off-shell formulation of the $\mathcal{N} = 8$ theory?
- Notice that we have an invariant on-shell formulation:
 - Let τ be a complex 1-form with only holomorphic dependence on η^i .
 - Assume that rank $\{\tau \wedge \mathrm{d}\tau\} = \mathrm{3}|\mathcal{N}$ and

$$\mathsf{ker}\{\tau \wedge \mathsf{d}\tau\} \cap \overline{\mathsf{ker}\{\tau \wedge \mathsf{d}\tau\}} = \{\mathbf{0}\}.$$

• Then define $T^{0,1}P := \ker\{\tau \land d\tau\}$ and apply the theorem.

• What about the full $\mathcal{N} = 8$ Einstein supergravity and the relation to twistor strings?

• ...

Thank you very much for your attention!

Martin Wolf $\mathcal{N} = 8$ Self-Dual Supergravity