

Random Effect Modelling with Mode Trees

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Work supported by

Framework: Generalized linear model with random effect

$$\mu_i \equiv E(y_i | z_i, \beta) = h(\eta_i) \equiv h(x_i' \beta + z_i).$$

The marginal likelihood can be approximated by a finite mixture (Laird, 1978)

$$L(\beta, g(z)) = \prod_{i=1}^n \int f(y_i | z_i, \beta) g(z_i) dz_i \approx \prod_{i=1}^n \left\{ \sum_{k=1}^K f(y_i | z_k, \beta) \pi_k \right\}$$

with mass points z_k and masses π_k .

Two well-known approaches:

- $Z \sim N(0, \sigma^2) \longrightarrow$ **Gaussian Quadrature** (Hinde, 1982)
- No distributional assumption about $Z \longrightarrow$ **Nonparametric maximum likelihood** (NPML, Aitkin, 1996).

A special case: Fitting finite Gaussian mixtures

Given: A set of observations y_1, \dots, y_n with

- $y_i \sim N(z_i, \sigma_i^2), i = 1, \dots, n$
- $z_i \sim Z$, where Z is left unspecified.

Random effect model without fixed term:

$$\mu_i = E(y_i | z_i) = z_i$$

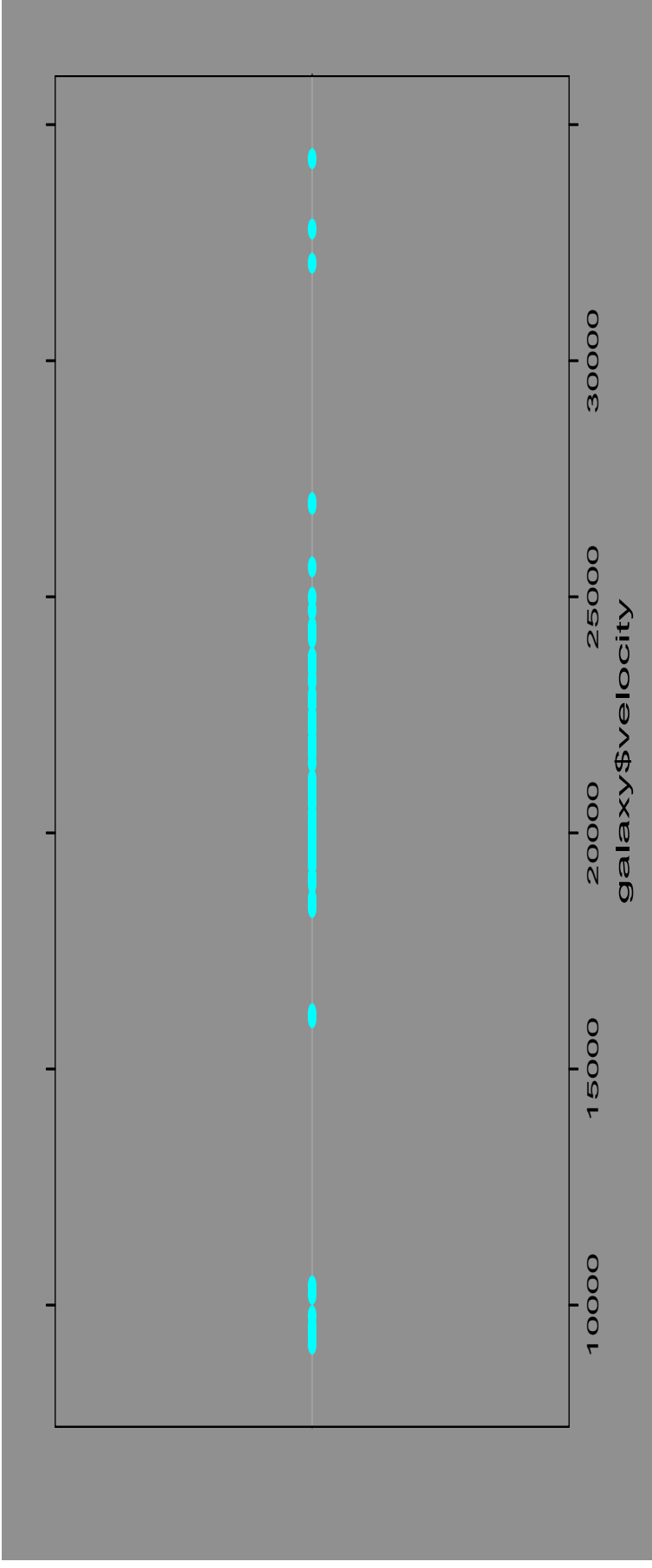
Mixture density:

$$f(y | (z_k, \pi_k, \sigma_k)_{k=1, \dots, K}) = \sum_{k=1}^K \pi_k f(y | z_k, \sigma_k^2)$$

where $f(y | z_k, \sigma_k^2)$ is a normal density with mean z_k and standard deviation σ_k .

Example: Galaxy Data

Recession velocities (in km/s) of 82 galaxies.



Aim: Estimate the **mass points** z_k , the **variances** σ_k^2 , the **masses** π_k , and the **number** of **components** K .

NPML Estimation

For fixed K , consider the log-likelihood

$$\ell = \sum_{i=1}^n \log \left\{ \sum_{k=1}^K \pi_k f(y_i | z_k, \sigma_k^2) \right\},$$

and calculate the score equations

$$\frac{\partial \ell}{\partial z_k} = 0, \quad \frac{\partial \ell - \lambda(\sum \pi_k - 1)}{\partial \pi_k} = 0, \quad \frac{\partial \ell}{\partial \sigma_k} = 0,$$

which turn out to be weighted versions of the single-distribution score equations.

\implies can be solved by standard EM algorithm:

Starting points Select starting values z_k^0 , π_k^0 , and σ_k^0 , $k = 1, \dots, K$.

E-Step Adjust weights given current parameter estimates.

M-Step Update parameter estimates.

Application on galaxy data

Set e.g. $K=5$:

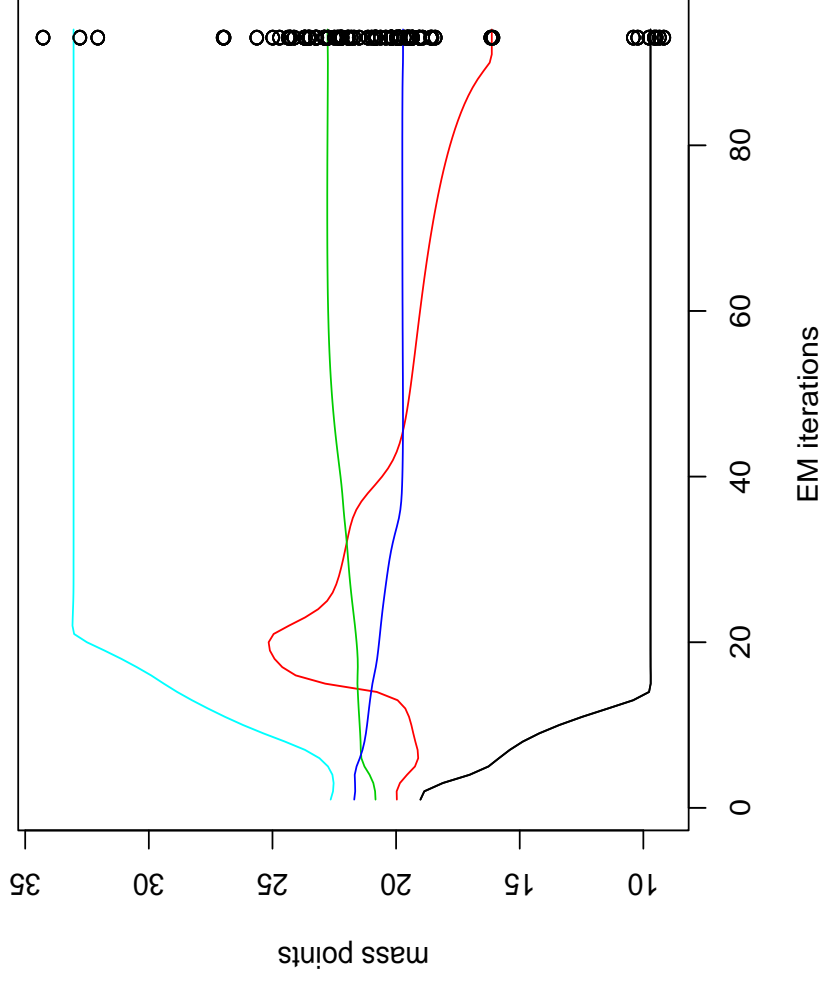
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Coefficients:
  MASS1  MASS2  MASS3  MASS4  MASS5
  9.71  16.13  22.78  19.72  33.04

Mixture proportions:
  0.085  0.024  0.512  0.342  0.037

Standard deviations:
  0.423  0.043  1.721  0.626  0.922

-2 log L:      380.9
```

EM Trajectories:



Properties of current NPML implementations (as in GLIM 4)

- The algorithm is simple, converges in every case, and is "impressively stable" (Aitkin, 1996).

- Results depend heavily on the choice of starting points z_k^0 , usually defined as

$$z_k^0 = \bar{y} + \text{tol} * \hat{\sigma} * g_k$$

where tol : scaling parameter, g_k : Gauss-Hermite mass points, $\hat{\sigma} = \frac{1}{n} \sum (y_i - \bar{y})^2$.

Finding the optimal solution requires a tedious grid search for tol .

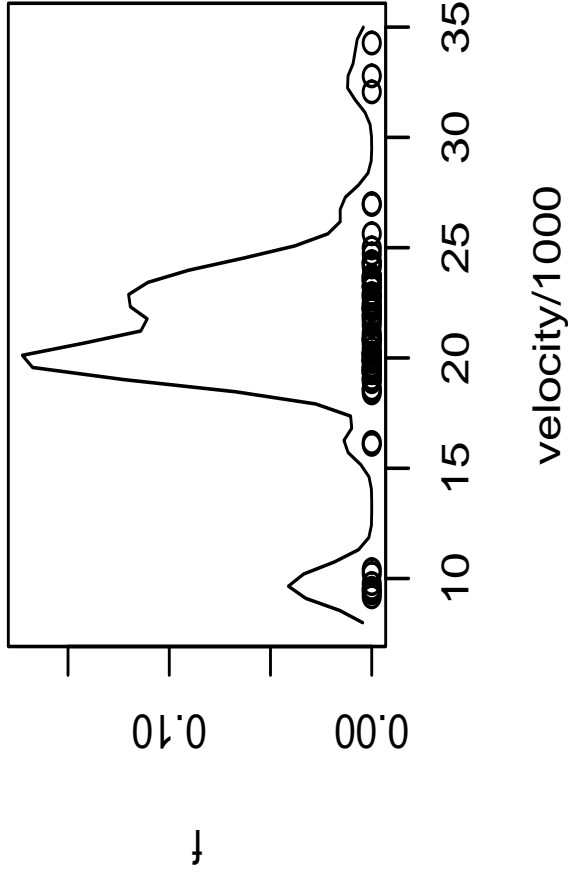
- The EM trajectories behave quite erratically in the first cycles, and tend to cross.
- There does not exist any automatic routine to select K . Richardson & Green, 1997:

"one of the things you do not know is the number of things that you do not know"

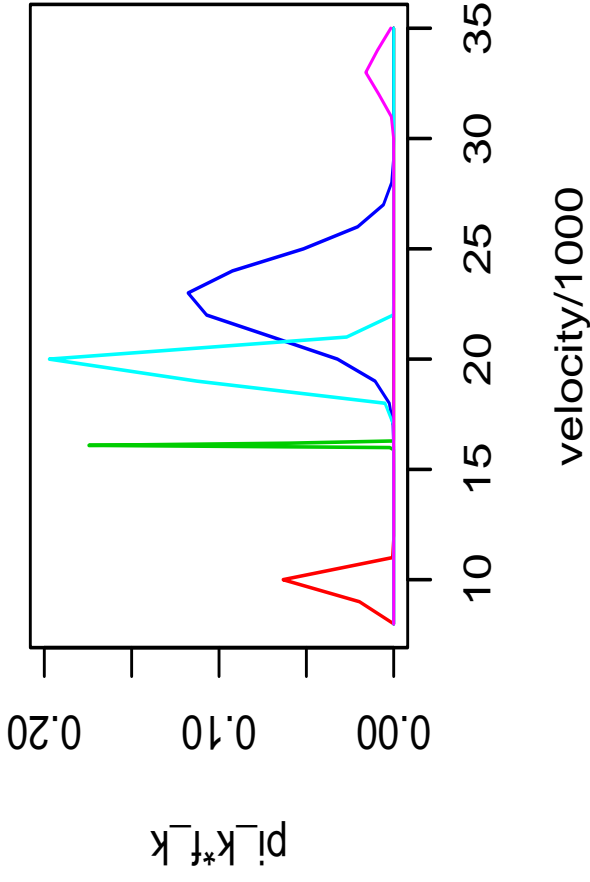
Exploiting the multimodal structure

Idea: Consider density estimate $\hat{f}(y, h) = \frac{1}{nh} \sum_{i=1}^n K\left(\frac{Y_i - y}{h}\right)$.

estimated density

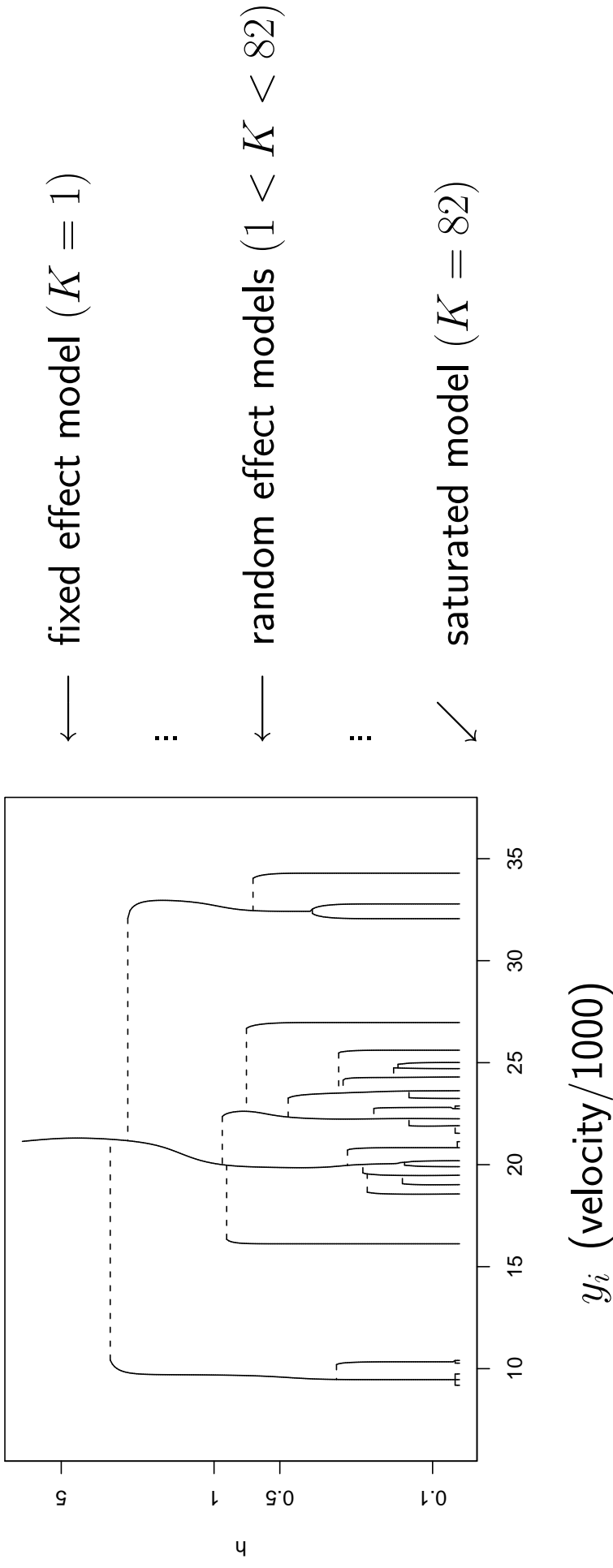


est. mixture components



Carreira-Perpiñan & Williams, 2003: The number of modes is a *lower bound* for the number of components of a Gaussian mixture. **But:** Number of modes depends on bandwidth h .

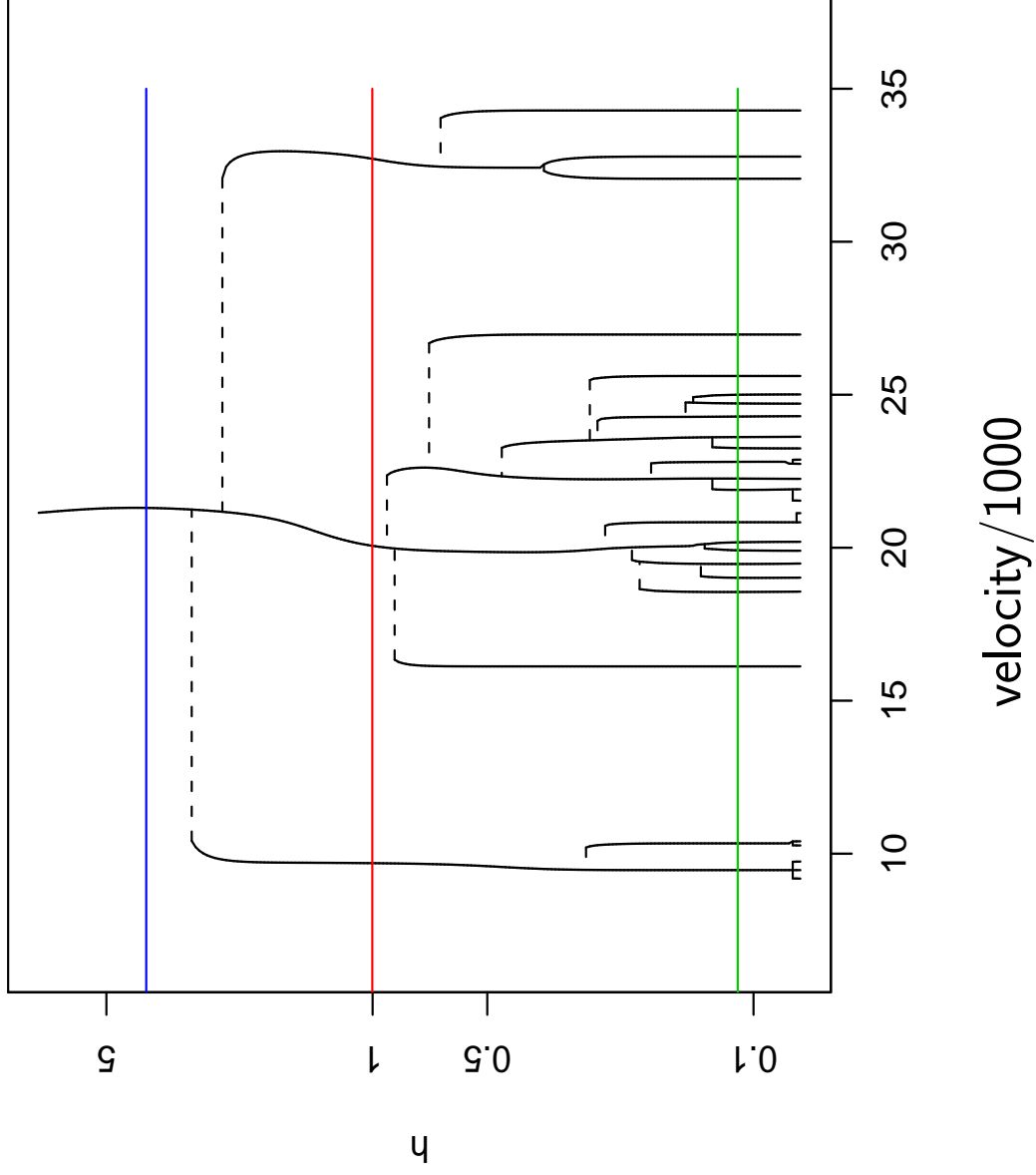
The mode tree (Minnotte & Scott, 1993)



More generally, applied on the 'residuals' $h^{-1}(y_i) - x_i'\hat{\beta}$ of a GLM:

”zoom into the random effect distribution”

Examples for bandwidth selection



Bandwidth selectors:

BCV (1 mode)

Silverman (3 modes)

AIC (22 modes)

Bandwidth selection in 2 steps

- Calculate Silverman's optimal bandwidth

$$h_{opt} = 0.9An^{-1/2},$$

where $A = \min\{\hat{\sigma}, IQR/1.34\}$

- From that bandwidth, climb down the mode tree until the next **critical bandwidth** (Silverman, 81)

$$h_{crit} = \inf\{h, \hat{f}(\cdot, h) \text{ has at most } k \text{ modes}\}$$

is reached.

Using h_{crit} , the mode tree gives

- an estimate for the number of modes, and thus for the number of components
- a very accurate estimate for the location of the mass points, which can be either used directly as mass point estimates \hat{z}_k , or as starting points z_k^0 for the EM algorithm.

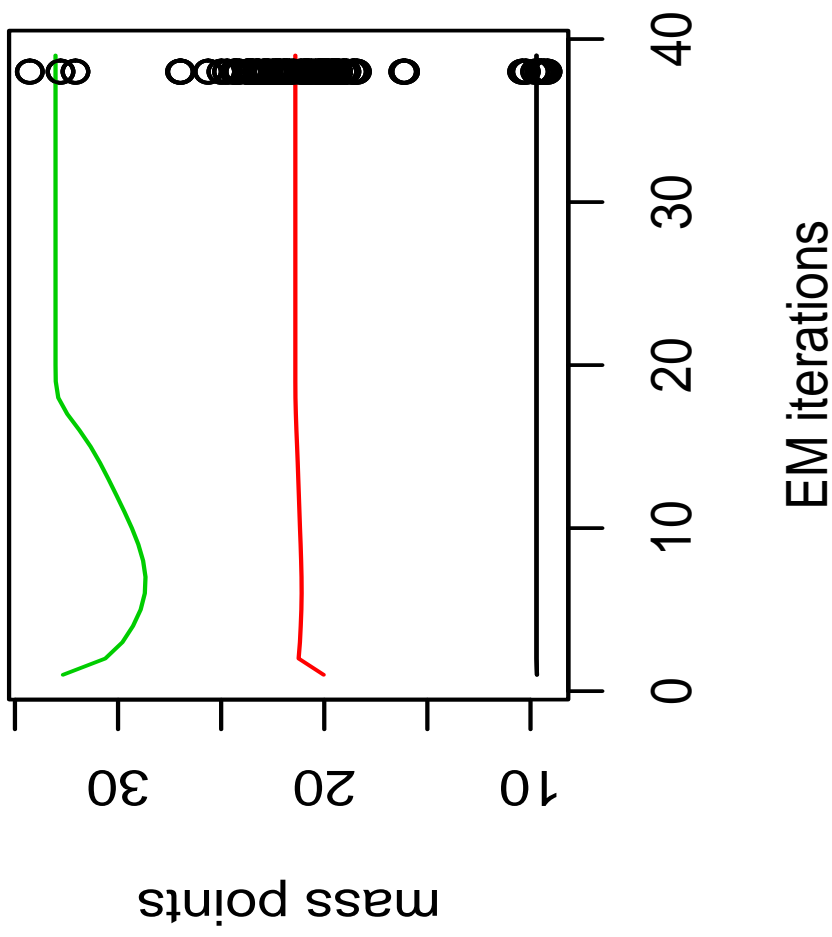
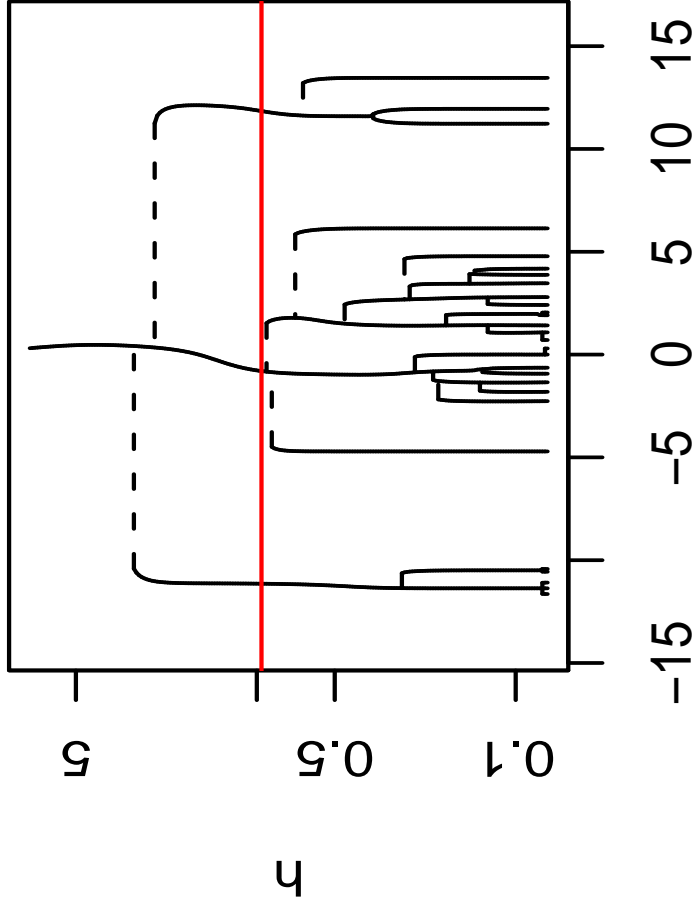
In the latter case, one has to **damp** the standard deviation in the initial loops of the EM algorithm via

$$d_j = 1 - (1 - \text{tol})^j, \quad (0 < \text{tol} \leq 1)$$

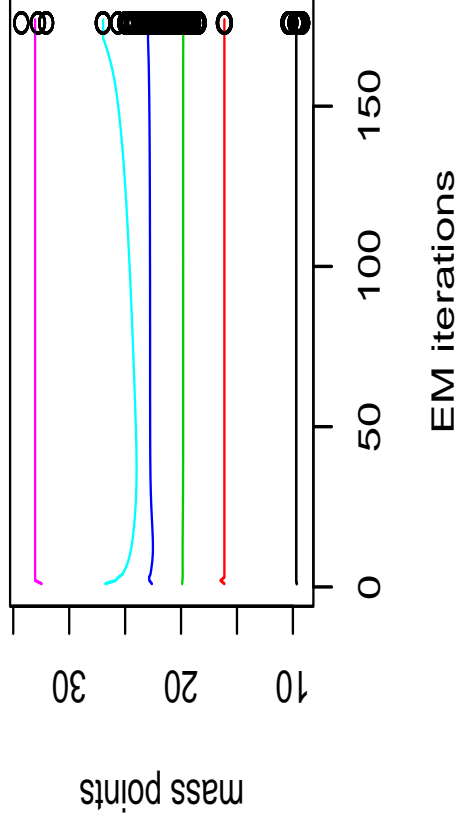
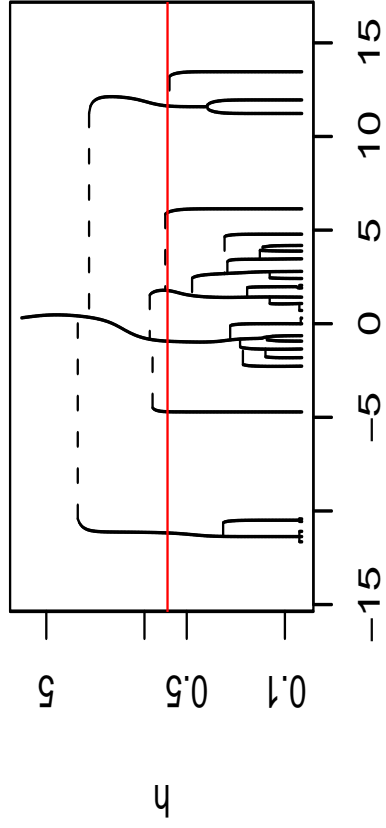
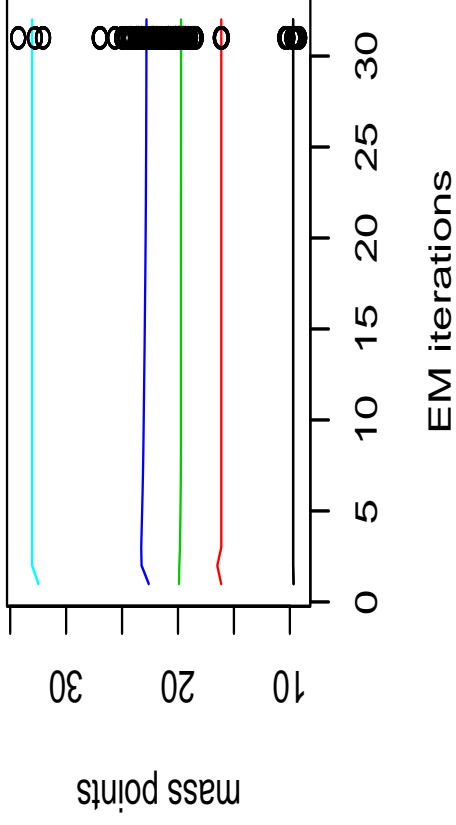
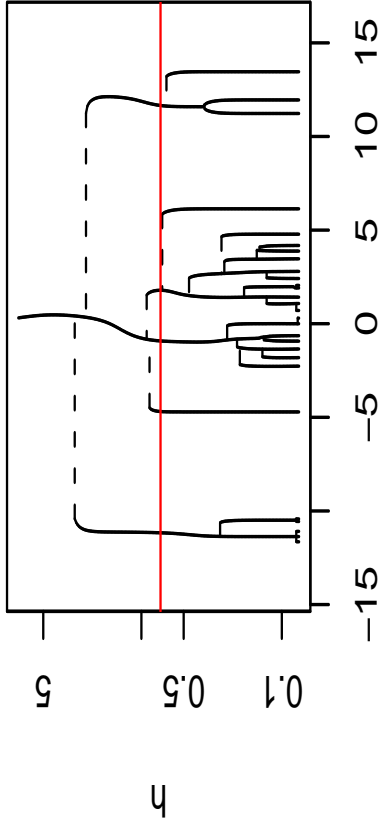
as otherwise the EM trajectories might be kicked out of the optimal solution immediately.

Application on galaxy data

Mode tree, **critical bandwidth**, and EM trajectories:



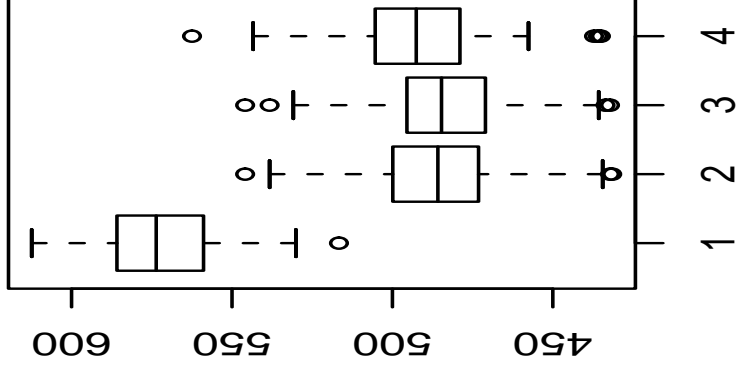
Climbing down the tree



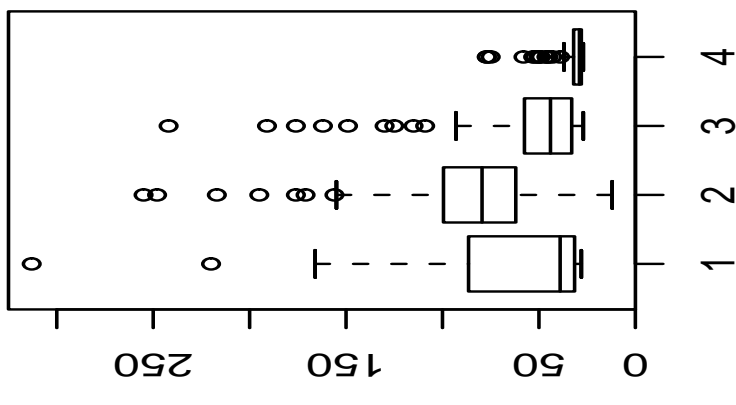
Simulation study: 100 samples from fitted mixture distribution ($\sigma = \sigma_k = const$).

Starting points z_k^0	Gauss-Hermite	Density Modes
Mass Points z_k	via EM algorithm	
Estimation Method	GQ	Mode Trees
	(1)	(2)
	NPML	NPML
	(3)	(4)

-2logL



EM iterations



Summary

- The mode tree is a useful instrument to assess visually the number of components, even if the multimodal structure cannot be seen in the data cloud itself.
- Mode trees together with a suitable bandwidth selector give a useful *recommendation* for the choice of K . However, this is no reliable automatic routine, as small variations in the bandwidth may change drastically the number of detected modes.
- Given K , mode trees give a set of estimated mass points, which are in many cases so accurate that one hardly needs EM at all!
- When using mode trees together with the proposed damping procedure, the sensitivity to tuning parameters (in particular tol) is drastically reduced.

Everything more general....

- Replace Gaussian by another exponential family distribution
- Set an appropriate link function
- Include explanatory variables
- Random coefficient models
- Variance component models

..... is being implemented in an R package {npml} (Einbeck, Darnell, & Hinde), see

www.nuigalway.ie/maths/je/npml.html

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