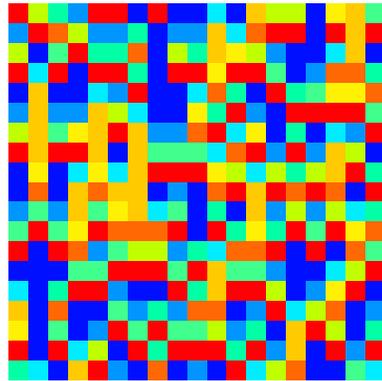




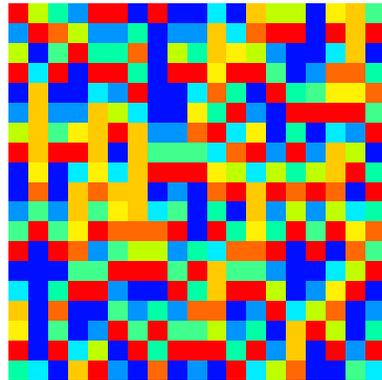
# STATISTICAL MECHANICS OF NONHYPERBOLIC COUPLED MAP LATTICES



Christian Beck  
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- 1 Introduction
- 2 Diffusively coupled Tchebyscheff maps
- 3 Observed scaling behaviour
- 4 Perturbative result for the invariant density
- 5 Physical relevance of CMLs in quantum field theory and cosmology
- 6 Summary



## 1 Introduction

large 1-dim lattices, lattice sites  $i$ . Dynamics given by

$$\Phi_{n+1}^i = (1 - a)T(\Phi_n^i) + \frac{a}{2}(T(\Phi_n^{i-1}) + T(\Phi_n^{i+1}))$$

$i$ : discrete spatial coordinate (periodic boundary conditions)

$n$ : discrete time

$a$ : coupling constant

$T$ : local map, e.g.  $T(\Phi) = 2\Phi^2 - 1$  (negative Ulam map)



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Colour coding:

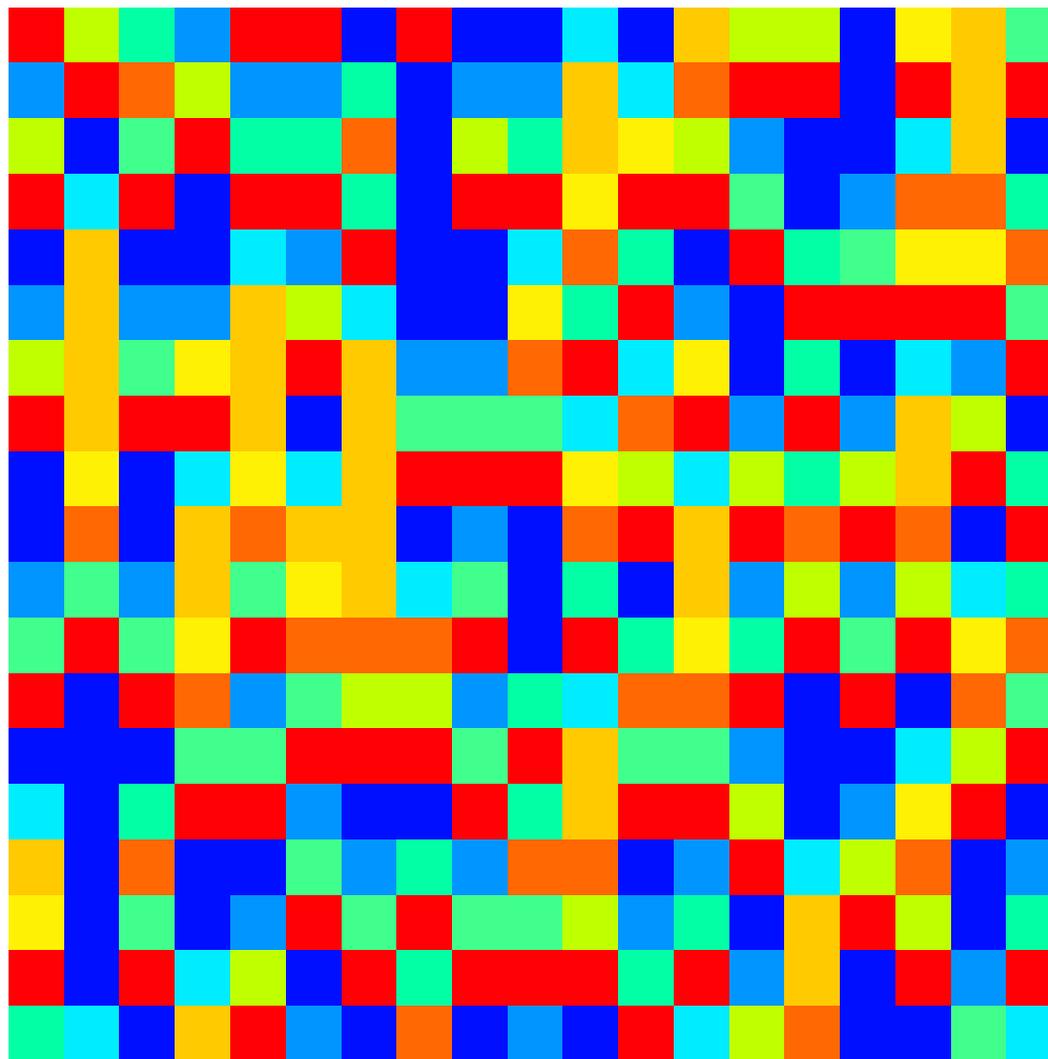


-1.0 -0.8 -0.6 -0.4 -0.2 0.00 +0.2 +0.4 +0.6 +0.8 +1.0

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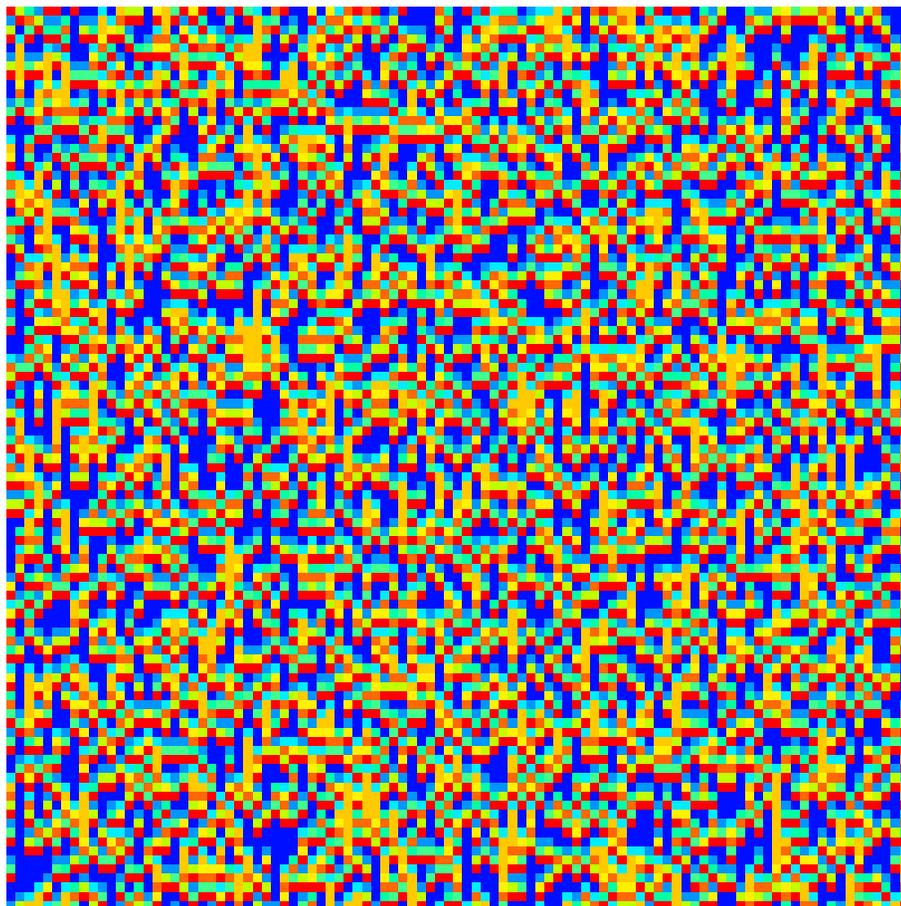
$i \rightarrow$

$n \downarrow$

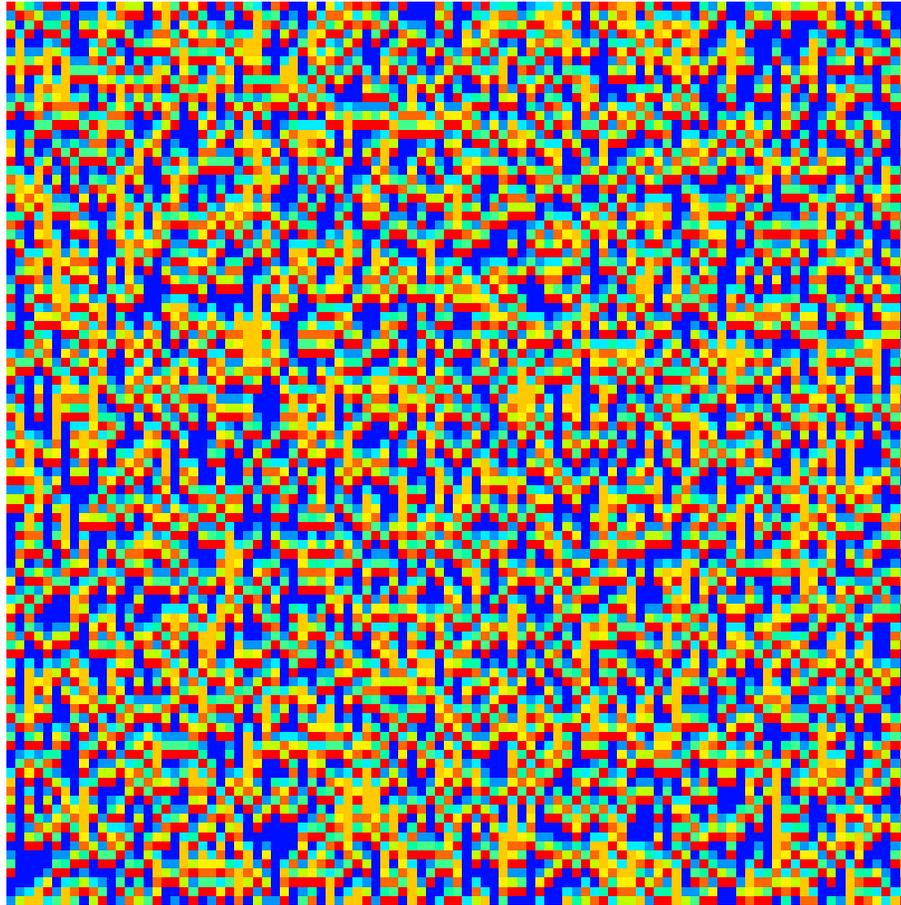


$a=0$

(larger lattice)



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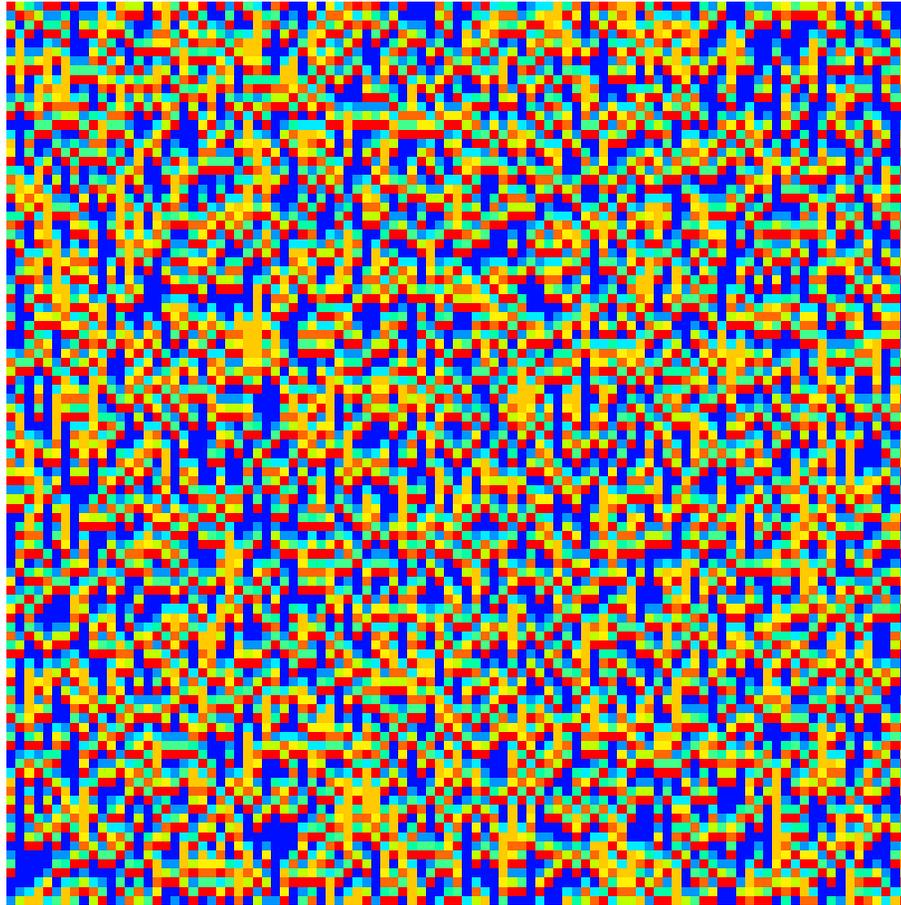


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Ulam map conjugated to tent map, iterates satisfy a **Central Limit Theorem** for a=0:

$$\frac{1}{\sqrt{M}} \sum_{n=1}^M \Phi_n^i \rightarrow \text{Gaussian} (M \rightarrow \infty)$$



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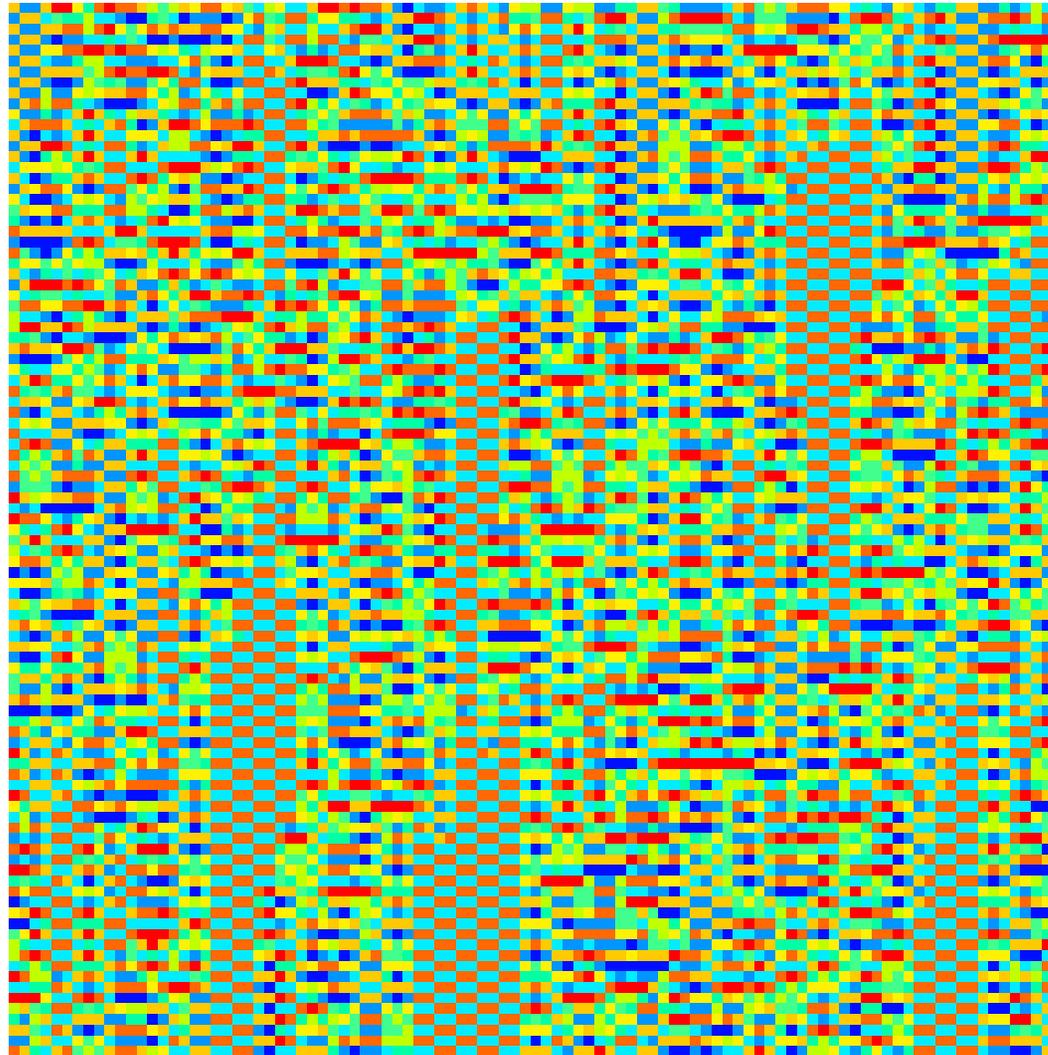
$$\frac{1}{\sqrt{M}} \sum_{n=1}^M \Phi_n^i \rightarrow \text{Gaussian} (M \rightarrow \infty)$$

But there are complicated higher-order correlations, see C.B., Nonlinearity 4, 1131 (1991)

$n$ -point functions  $\langle \Phi_{n_1}^i \Phi_{n_2}^i \dots \Phi_{n_r}^i \rangle$  do **not** factorize, even for  $a = 0$ .

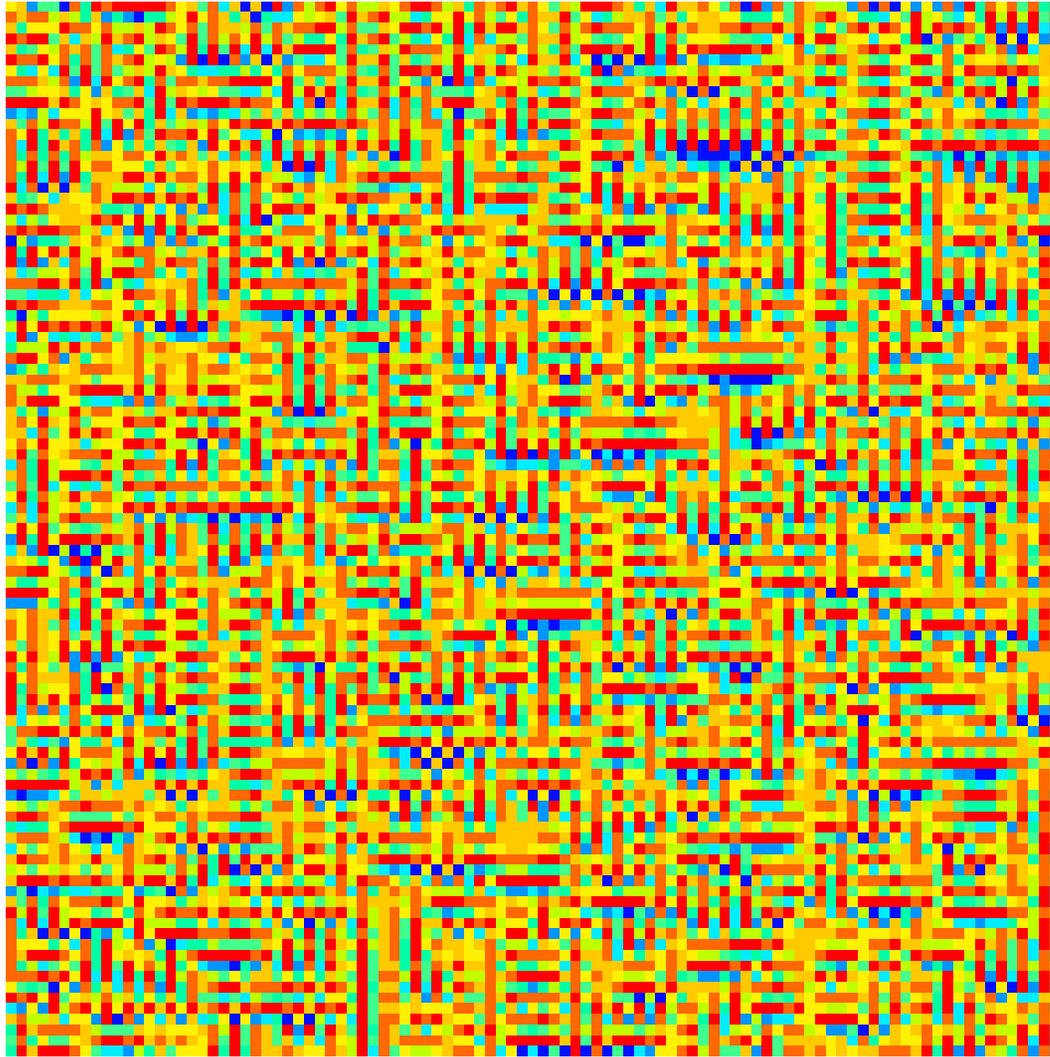
These correlations are 'scaled away' for  $M \rightarrow \infty$ .

$$\Phi_{n+1}^i = (1 - a)T(\Phi_n^i) + \frac{a}{2}(T(\phi_n^{i-1}) + T(\Phi_n^{i+1}))$$

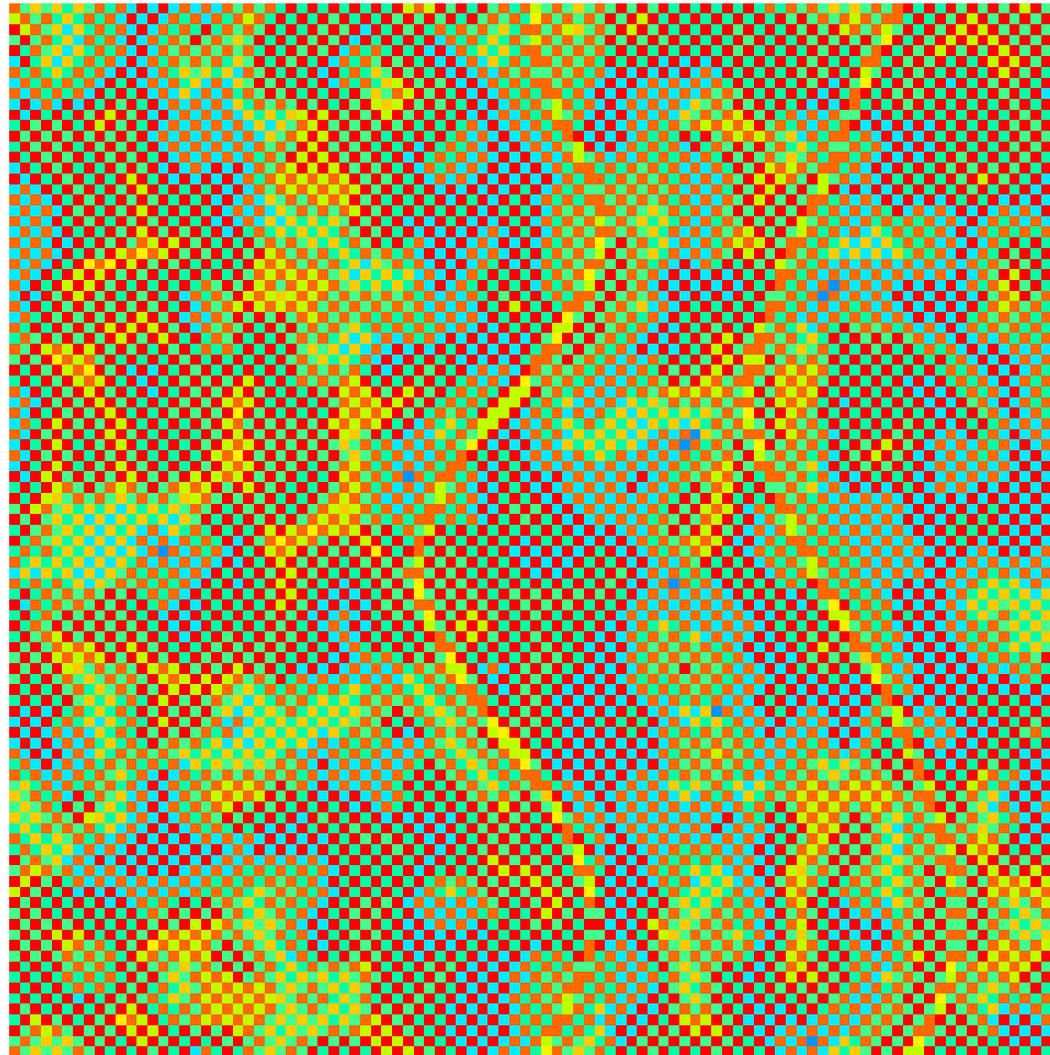


$a = 0.375$

$a = 1$



$a=0.5$ , 2-dim lattice



(snapshot at fixed time  $n$ )



## 2 Diffusively coupled Tschebyscheff maps

Lots of results for CMLs consisting of **hyperbolic** (uniformly expanding) maps

(hyperbolicity in 1-d case:  $|slope| > 1$ ) (e.g. work by Keller, Kuenzle, Jarvenpaa, Baladi, Rugh, MacKay, Bunimovich, Just, Pesin,...)

but much less is known for **nonhyperbolic** situations, though some promising steps have been made (Chaté, Torcini, Ruffo, ...)

We are particularly interested in cases where the local map exhibits strongest possible chaotic behaviour and **small** coupling, e.g. Tchebyscheff maps  $T_N$  of  $N$ -th order:

$$T_2(\Phi) = 2\Phi^2 - 1 \quad (1)$$

$$T_3(\Phi) = 4\Phi^3 - 3\Phi \quad (2)$$

$$\dots = \dots \quad (3)$$

$$T_N(\Phi) = \cos(N \arccos \Phi) \quad (4)$$

conjugated to a Bernoulli shift of  $N$  symbols (generalized tent maps with  $|slope| = N$  having  $\sim N/2$  maxima).



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This conjugacy is 'destroyed' for finite coupling  $a > 0$  in the CML.



### 3 Observed scaling behavior of invariant density

Single Tchebyscheff map: Invariant density given by  $\rho_0(x) = \frac{1}{\pi \sqrt{1-x^2}}$ .

CML with  $a = 0$ : The invariant density for all  $M$  lattice sites is (of course) given by

$$\rho_0(\Phi^1, \Phi^2, \dots, \Phi^M) = \prod_{i=1}^M \frac{1}{\pi \sqrt{1 - (\Phi^i)^2}} \quad (5)$$



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Note that this is like a generalized canonical ensemble (product of  $q$ -exponentials) in nonextensive statistical mechanics with  $q = 3$ , energy  $\epsilon = \frac{1}{2}\Phi^2$ ,  $\beta = 1$ .

(recall  $e_q(x) := (1 + (q - 1)x)^{-\frac{1}{q-1}}$ , hence  $\rho_0(\phi) = \frac{1}{\pi}e_q^{-\beta\epsilon}$  (formal analogy only))



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For finite  $a > 0$  the density changes and is not a product of single-site densities any more.

Still one can define the **1-point density**  $\rho_a(\Phi)$  at each lattice site as a marginal density (integrating the joint density over all but one lattice site).

We are interested in averages of arbitrary single-site test functions  $h(\Phi^i)$ :

$$\langle h(\Phi) \rangle_a = \lim_{M \rightarrow \infty, J \rightarrow \infty} \frac{1}{MJ} \sum_{n=1}^M \sum_{i=1}^J h(\Phi_n^i). \quad (6)$$

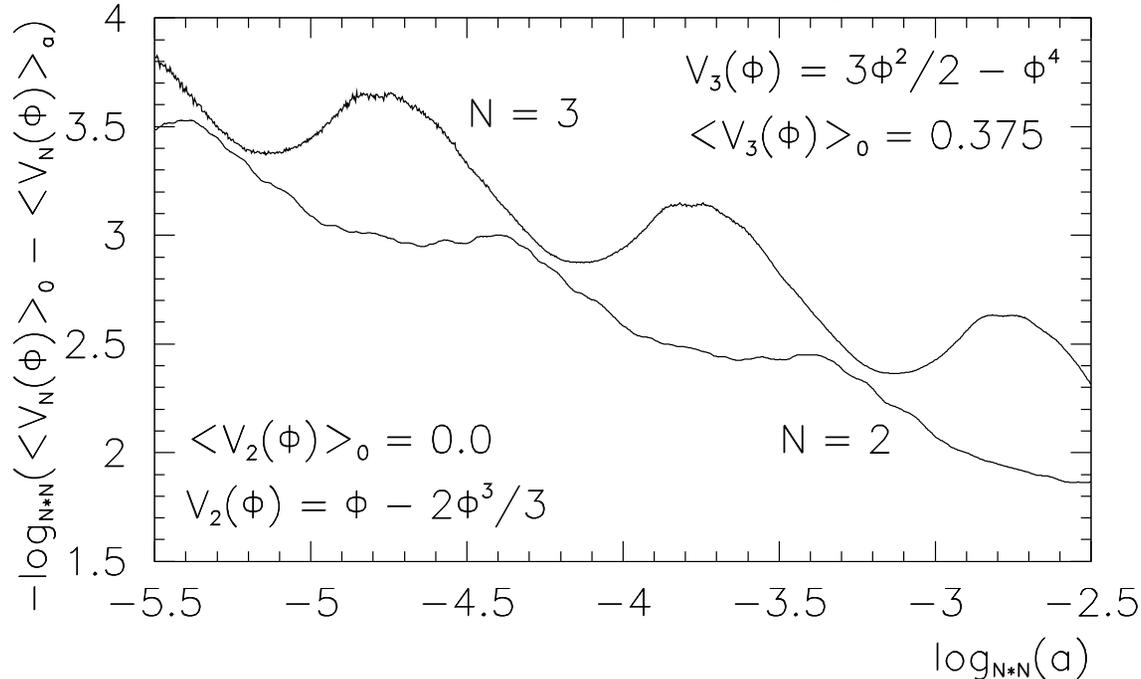
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For  $a \rightarrow 0$  one numerically observes the **scaling behaviour**

$$\langle h(\Phi) \rangle_a - \langle h(\Phi) \rangle_0 = \sqrt{a} \cdot F^{(N)}(\log a) \quad (7)$$

where  $F^{(N)}$  is a periodic function of  $\log a$  with period  $\log N^2$ .



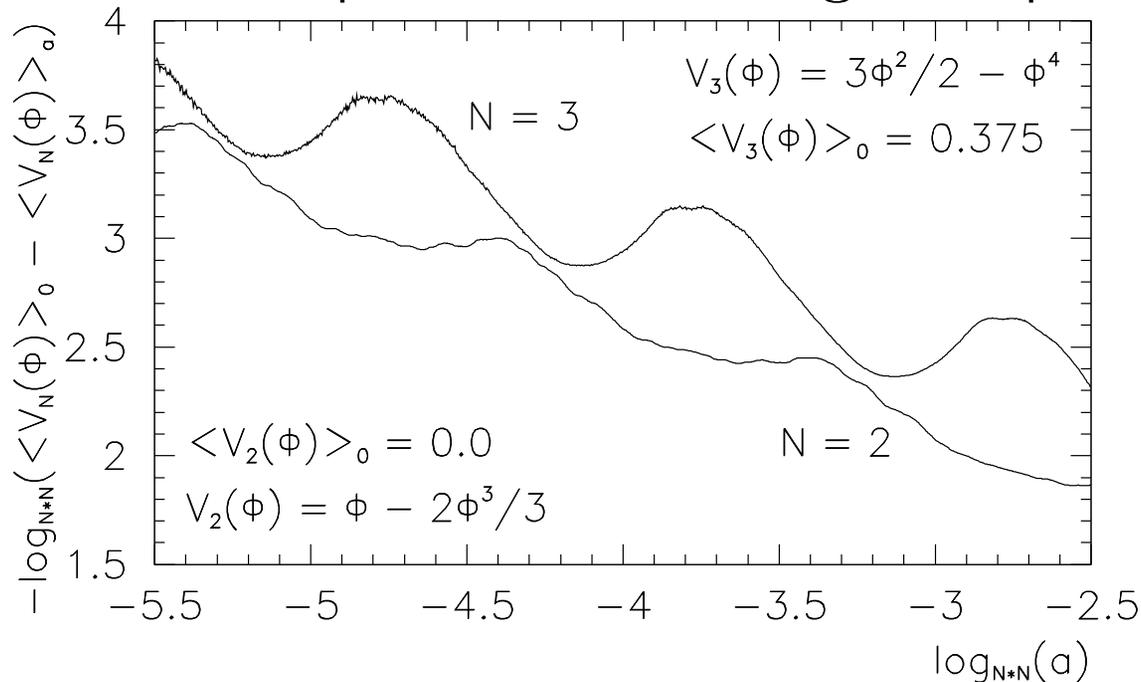
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chosen test functions in the above plot :

$$h(\Phi) = \Phi - \frac{2}{3}\Phi^3 \quad (N = 2) , \quad h(\Phi) = \frac{3}{2}\Phi^2 - \Phi^4 \quad (N = 3)$$

Not only scaling behaviour in the **parameter space** but also in the **phase space**:

Near the left edge of the interval  $[-1, 1]$  we may write  $\Phi = ay - 1$  and observe the scaling behaviour

$$\rho_a(ay - 1) = a^{-1/2}g(y) \quad (8)$$

where the function  $g$  is independent of  $a$  for small  $a$ .

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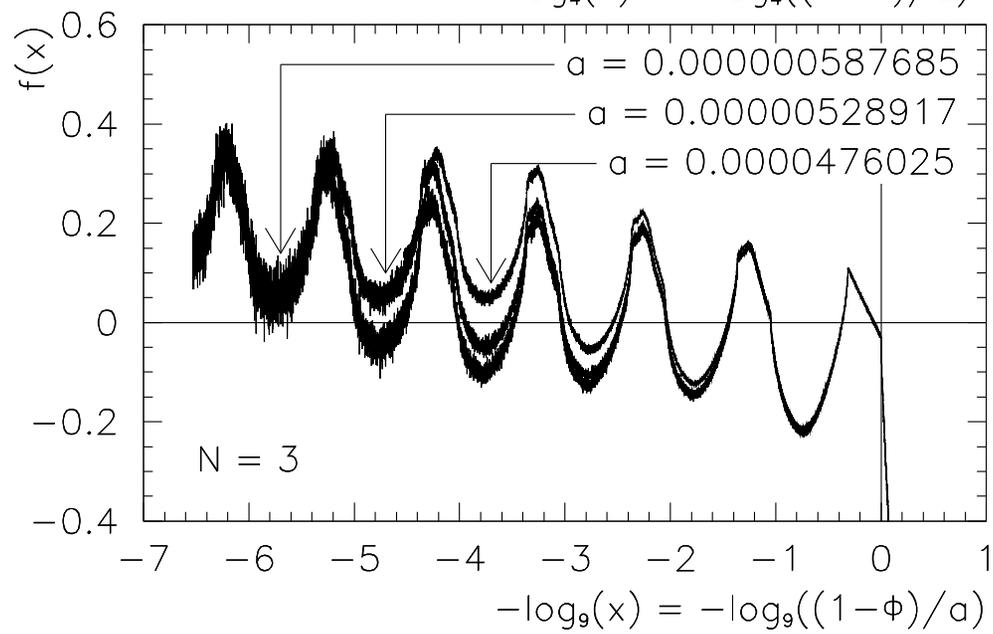
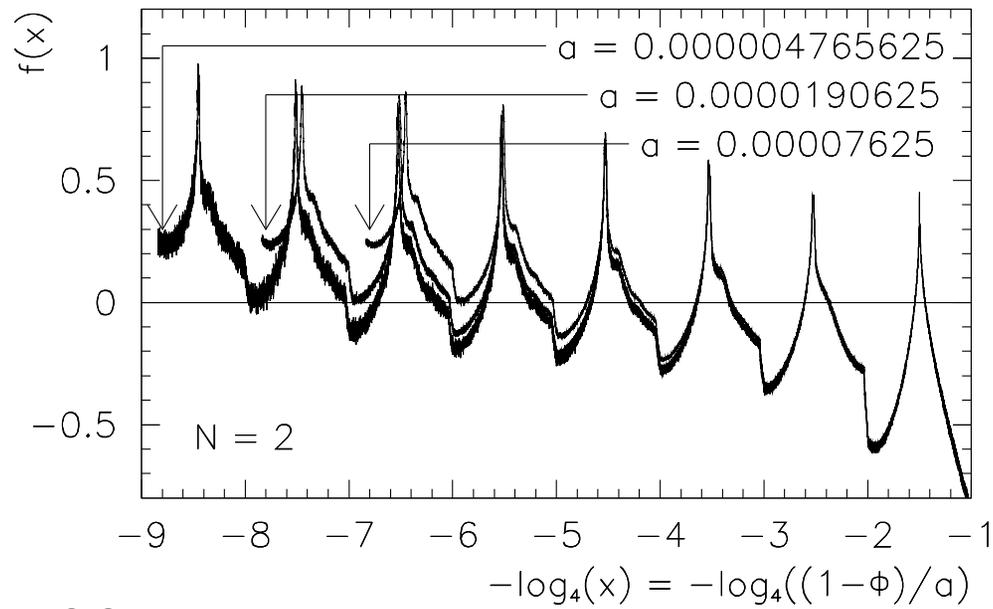
At the right edge, writing  $\Phi = 1 - ax$ , one observes

$$\rho_a(1 - ax) = \rho_0(1 - ax) + \frac{1}{2}a^{-1/2}x^{-1}f(x) \quad (9)$$

where  $f$  is independent of  $a$  for small  $a$ . Moreover,  $f$  exhibits log-periodic oscillations

$$f(N^2x) = f(x) \quad (10)$$

over a large region of the phase space.





#### 4 Perturbative results for the invariant 1-point density

Final result of a longer calculation ( $N = 2$ ):

Left edge: In leading order in  $a$

$$\rho_a(-1 + ay) = \frac{1}{\pi\sqrt{2a}} \int_{1-y}^1 \frac{\rho_{00}(z) dz}{\sqrt{y-1+z}} \quad (11)$$

$$\rho_{00}(z) = \frac{2}{\pi^2} K(\sqrt{1-z^2}) \theta(1-z^2) \quad (12)$$

where  $K(x)$  is the complete elliptic integral of the first kind.

right edge:

$$\rho_a(1 - ax) = \sum_{p=1}^{\infty} \rho_a^{(p)}(1 - ax) \quad (13)$$

where

$$\rho_a^{(p)}(1 - ax) = \frac{1}{4^p \pi \sqrt{2a}} \int \frac{\rho_0(\phi_+) d\phi_+ \rho_0(\phi_-) d\phi_-}{\sqrt{x/4^p + r_2^p(\phi_+) + r_2^p(\phi_-)}}. \quad (14)$$

Here the function  $r_2^p(\phi)$  is defined as follows:

$$r_2^p(\phi) = \frac{1}{2} \sum_{q=0}^p \frac{T_{2q}(\phi) - 1}{2^{2q}}. \quad (15)$$

Limits of the two integrations in Eq. (14) given by the condition that  $|\phi_{\pm}| \leq 1$  and that the argument of the square root should always be positive.

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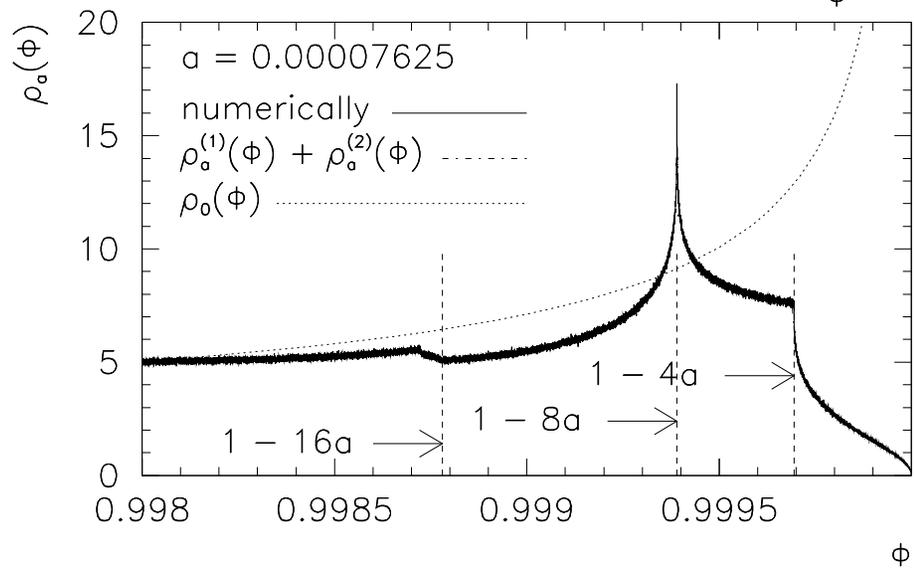
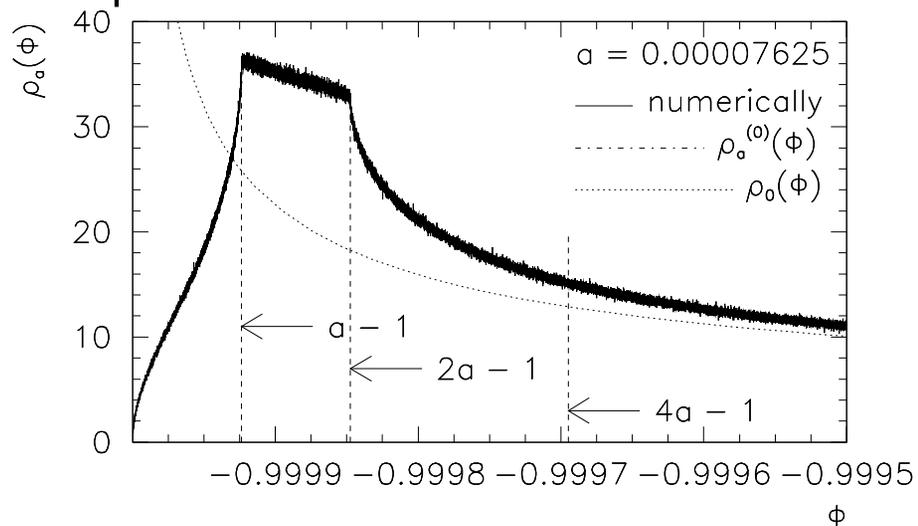
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Techniques: Start from perturbed 1-dimensional map, apply Perron-Frobenius and convolution techniques, iterate result

S. Groote, C.B., nlin.CD/0603397 (2006)

# Comparison with numerics:



Along similar lines, we obtain for  $N = 3$

$$\rho_a^{(p)}(1 - ax) = \frac{2}{9^p 3\pi \sqrt{2a}} \int \frac{\rho_0(\phi_+) d\phi_+ \rho_0(\phi_-) d\phi_-}{\sqrt{x/9^p + r_3^p(\phi_+) + r_3^p(\phi_-)}}, \quad (16)$$

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The density is symmetric, i.e.

$$\rho_a^{(p)}(ax - 1) = \rho_a^{(p)}(1 - ax). \quad (18)$$

For general  $N$ :

$$\rho_a^{(p)}(1 - ax) \sim \frac{1}{\sqrt{a}} \int \frac{\rho_0(\phi_+) d\phi_+ \rho_0(\phi_-) d\phi_-}{\sqrt{x/N^{2p} + r_N^p(\phi_+) + r_N^p(\phi_-)}}, \quad (19)$$

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Using these results, one can prove the existence of the log-periodic oscillations both in phase and parameter space.



## 5 Physical applications for nonhyperbolic CMLs in quantum field theories and cosmology

How can chaotic coupled map lattices be relevant in quantum field theories?



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[Stochastic Quantization.](#)

Consider [classical](#) field described by an action  $S[\varphi]$ . Classical field equation:

$$\frac{\delta S}{\delta \varphi} = 0 \quad (21)$$

meaning: Action has an extremum.



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How can chaotic coupled map lattices be relevant in quantum field theories?

**Stochastic Quantization.**

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$$\frac{\delta S}{\delta \varphi} = 0 \quad (21)$$

meaning: Action has an extremum.

Parisi-Wu (1981): Obtain **2nd quantized equation** of motion by considering a Langevin equation in **fictitious time  $s$** :

$$\frac{\partial}{\partial s} \varphi(x, s) = -\frac{\delta S}{\delta \varphi}(x, s) + L(x, s) \quad (22)$$

$x = (x^1, x^2, x^3, x^4) = x^\mu$  point in Euclidean space-time

$x^4 = t$  physical time

$L(x, s)$  spatio-temporal Gaussian white noise

$$\langle L(x, s) \rangle = 0 \quad (23)$$

$$\langle L(x, s)L(x', s') \rangle = 2\delta(x - x')\delta(s - s') \quad (24)$$

Parisi and Wu: Quantum mechanical expectations = expectations of Langevin process for  $s \rightarrow \infty$ .

Example:  $\varphi^4$ -theory

Action:

$$S[\varphi] = \int d^4x \left( \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi - \frac{1}{2} m^2 \varphi^2 - \frac{\lambda}{4} \varphi^4 \right) \quad (25)$$

Classical field equation:

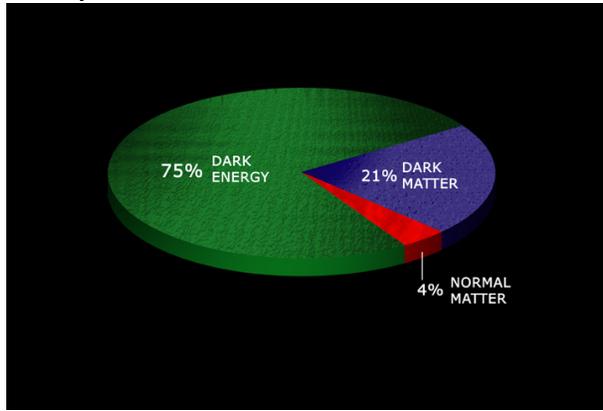
$$(-\partial^2 + m^2)\varphi(x) + \lambda\varphi^3(x) = 0 \quad (26)$$

2nd quantized version:

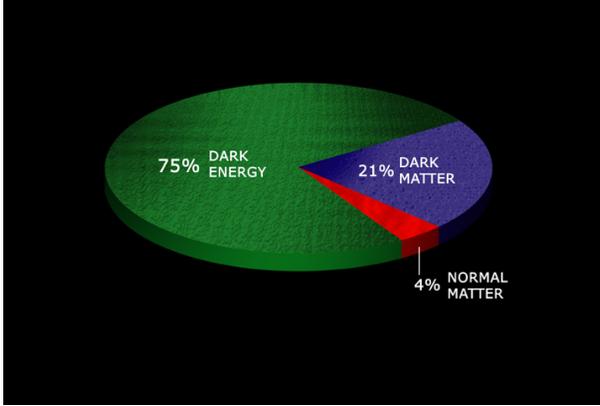
$$\frac{\partial}{\partial s} \varphi(x, s) = (\partial^2 - m^2)\varphi(x, s) - \lambda\varphi^3(x, s) + L(x, s) \quad (27)$$

Now construct a [chaotic dark energy model](#) based on a stochastically quantized scalar field (C. B., Phys. Rev. D 69, 123515 (2004))

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Quantized scalar field  $\varphi$  in Robertson-Walker metric:

$$\frac{\partial}{\partial s}\varphi = \ddot{\varphi} + 3H\dot{\varphi} + V'(\varphi) + L(s, t), \quad (28)$$

where  $H$  is the Hubble parameter,  $V$  is the potential under consideration and  $L(s, t)$  is Gaussian white noise  $t$  physical time,  $s$  fictitious time. Discretize

$$s = n\tau \quad (29)$$

$$t = i\delta \quad (30)$$

$\tau$ : fictitious time lattice constant,  $\delta$ : physical time lattice constant. We obtain

$$\frac{\varphi_{n+1}^i - \varphi_n^i}{\tau} = \frac{1}{\delta^2}(\varphi_n^{i+1} - 2\varphi_n^i + \varphi_n^{i-1}) + 3\frac{H}{\delta}(\varphi_n^i - \varphi_n^{i-1}) + V'(\varphi_n^i) + \text{noise} \quad (31)$$

This can be written as the following recurrence relation for the field  $\varphi_n^i$

$$\varphi_{n+1}^i = (1-\alpha) \left\{ \varphi_n^i + \frac{\tau}{1-\alpha} V'(\varphi_n^i) \right\} + 3 \frac{H\tau}{\delta} (\varphi_n^i - \varphi_n^{i-1}) + \frac{\alpha}{2} (\varphi_n^{i+1} + \varphi_n^{i-1}) + \tau \cdot n \cdot \varphi_n^i \quad (32)$$

where a dimensionless coupling constant  $\alpha$  is introduced as

$$\alpha := \frac{2\tau}{\delta^2}. \quad (33)$$

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Introduce dimensionless field variable  $\Phi_n^i$  by writing  $\varphi_n^i = \Phi_n^i p_{max}$ , where  $p_{max}$  is some (so far) arbitrary energy scale.  $\implies$

$$\Phi_{n+1}^i = (1-\alpha) T(\Phi_n^i) + \frac{3}{2} H \delta \alpha (\Phi_n^i - \Phi_n^{i-1}) + \frac{\alpha}{2} (\Phi_n^{i+1} + \Phi_n^{i-1}) + \tau \cdot \text{noise}, \quad (34)$$

where the local map  $T$  is given by

$$T(\Phi) = \Phi + \frac{\tau}{p_{max}(1-\alpha)} V'(p_{max}\Phi). \quad (35)$$

Note that a symmetric diffusively coupled map lattice (Kaneko 1984)

$$\Phi_{n+1}^i = (1 - \alpha)T(\Phi_n^i) + \frac{\alpha}{2}(\Phi_n^{i+1} + \Phi_n^{i-1}) + \tau \cdot \textit{noise} \quad (36)$$

is obtained if  $H\delta \ll 1$ , equivalent to

$$\delta \ll H^{-1} \quad (37)$$

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The main result of our consideration is that iteration of a coupled map lattice of the form (36) with a given map  $T$  has physical meaning: It means that one is considering the second-quantized dynamics of a self-interacting real scalar field  $\varphi$  with a force  $V'$  given by

$$V'(\varphi) = \frac{1 - \alpha}{\tau} \left\{ -\varphi + p_{max} T \left( \frac{\varphi}{p_{max}} \right) \right\}. \quad (38)$$

Integration yields

$$V(\varphi) = \frac{1 - \alpha}{\tau} \left\{ -\frac{1}{2}\varphi^2 + p_{max} \int d\varphi T \left( \frac{\varphi}{p_{max}} \right) \right\} + \text{const.} \quad (39)$$

In terms of the dimensionless field  $\Phi$  this can be written as

$$V(\varphi) = \frac{1 - \alpha}{\tau} p_{max}^2 \left\{ -\frac{1}{2}\Phi^2 + \int d\Phi T(\Phi) \right\} + \text{const.} \quad (40)$$

Lattice constant  $\tau$  should be small, in order to approximate the continuum theory, which is ordinary quantum field theory. **Typical choice  $\tau \sim 1/m_{Pl}^2$ .**

Distinguished example of a  $\varphi^4$ -theory generating **strongest possible chaotic behaviour**:

$$\Phi_{n+1} = T_{-3}(\Phi_n) = -4\Phi_n^3 + 3\Phi_n \quad (41)$$

on the interval  $\Phi \in [-1, 1]$ .  $T_{-3}$  is the negative third-order Tchebyscheff map, a standard example of a map exhibiting strongly chaotic behaviour. It is conjugated to a **Bernoulli shift**. The corresponding potential is given by

$$V_{-3}(\varphi) = \frac{1 - \alpha}{\tau} \left\{ \varphi^2 - \frac{1}{p_{max}^2} \varphi^4 \right\} + \text{const}, \quad (42)$$

or, in terms of the dimensionless field  $\Phi$ ,

$$V_{-3}(\varphi) = \frac{1 - \alpha}{\tau} p_{max}^2 (\Phi^2 - \Phi^4) + \text{const}. \quad (43)$$

We obtain by second quantization a field  $\varphi$  that rapidly fluctuates in fictitious time on some finite interval, provided that initially  $\varphi_0 \in [-p_{max}, p_{max}]$ .

Of physical relevance are the **expectations** of suitable observables with respect to the ergodic chaotic dynamics. For example, the expectation  $\langle V_{-3}(\varphi) \rangle$  of the potential is a possible candidate for vacuum energy in our universe. One obtains

$$\langle V_{-3}(\varphi) \rangle = \frac{1 - \alpha}{\tau} p_{max}^2 (\langle \Phi^2 \rangle - \langle \Phi^4 \rangle) + \text{const.} \quad (44)$$

For uncoupled Tchebyscheff maps ( $\alpha = 0$ ), expectations of any observable  $\mathbf{A}$  can be evaluated as the ergodic average

$$\langle \mathbf{A} \rangle = \int_{-1}^{+1} \mathbf{A}(\Phi) d\mu(\Phi), \quad (45)$$

with the natural invariant measure being given by

$$d\mu(\Phi) = \frac{d\Phi}{\pi \sqrt{1 - \Phi^2}} \quad (46)$$

From eq. (46) one obtains  $\langle \Phi^2 \rangle = \frac{1}{2}$  and  $\langle \Phi^4 \rangle = \frac{3}{8}$ , thus

$$\langle V_{-3}(\varphi) \rangle = \frac{1}{8} \frac{p_{max}^2}{\tau} + \text{const.} \quad (47)$$

Alternatively, we may consider the positive Tchebyscheff map  $T_3(\Phi) = 4\Phi^3 - 3\Phi$ . This basically exhibits the same dynamics as  $T_{-3}$ , up to a sign. Repeating the same calculation we obtain

$$V_3(\varphi) = \frac{1 - \alpha}{\tau} \left\{ -2\varphi^2 + \frac{1}{p_{max}^2} \varphi^4 \right\} + \mathit{const} \quad (48)$$

and

$$V_3(\varphi) = \frac{1 - \alpha}{\tau} p_{max}^2 (-2\Phi^2 + \Phi^4). \quad (49)$$

For the expectation of the vacuum energy one gets

$$\langle V_3(\varphi) \rangle = \frac{1 - \alpha}{\tau} p_{max}^2 (-2\langle \Phi^2 \rangle + \langle \Phi^4 \rangle) + \mathit{const}, \quad (50)$$

which for  $\alpha = 0$  reduces to

$$\langle V_3(\varphi) \rangle = -\frac{5p_{max}^2}{8\tau} + \mathit{const}. \quad (51)$$

Symmetry considerations between  $T_{-3}$  and  $T_3$  suggest to take the additive constant  $\mathit{const}$  as

$$\mathit{const} = +\frac{1 - \alpha}{\tau} p_{max}^2 \frac{1}{2} \langle \Phi^2 \rangle. \quad (52)$$

One obtains the fully symmetric equation

$$\langle V_{\pm 3}(\varphi) \rangle = \pm \frac{1 - \alpha}{\tau} p_{max}^2 \left\{ -\frac{3}{2} \langle \Phi^2 \rangle + \langle \Phi^4 \rangle \right\}, \quad (53)$$

which for  $\alpha \rightarrow 0$  reduces to

$$\langle V_{\pm 3}(\varphi) \rangle = \pm \frac{p_{max}^2}{\tau} \left( -\frac{3}{8} \right). \quad (54)$$

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The simplest model for dark energy in the universe, as generated by a chaotic  $\varphi^4$ -theory, would be to identify  $\frac{3}{8} p_{max}^2 / \tau = \rho_{\Lambda}$ , the constant vacuum energy density corresponding to a classical cosmological constant  $\Lambda$ , which stays constant during the expansion of the universe.

For a more sophisticated model (including late-time symmetry breaking due to structure formation, and tracking behaviour in the early universe) see C.B., Phys. Rev. D 69, 123515 (2004)

## Some interesting aspects of this model

- Vacuum fluctuations underlying dark energy are produced by a **deterministic chaotic noise field** (a CML) evolving in fictitious time
- Field (almost) conjugated to a Bernoulli shift today. Dynamics given by a CML of diffusively coupled 3rd-order Tchebyscheff map, coupling  $a \sim (m/m_{Pl})^2 \sim 10^{-50}$ . In the very early universe, coupling can be significantly larger.

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- Vacuum fluctuations underlying dark energy are produced by a **deterministic chaotic noise field** (a CML) evolving in fictitious time
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- Could these chaotic noise fluctuations help to 'derive' ordinary statistical mechanics, by coupling them as a small **Langevin-like noise term** to ordinary matter?
- Could these or similar types of fluctuations produce **measurable effects in laboratory experiments** (C.B., M.C.Mackey, Phys. Lett. B 605, 295 (2005))?



## 6 Summary

- Behaviour of **nonhyperbolic** CMLs much more complicated than that of hyperbolic ones.
- Scaling with  $\sqrt{a}$ , **logperiodic oscillations**, ...
- **Analytical** perturbative treatment possible for diffusively coupled Tchebyscheff maps of  $N$ -th order  
S. Groote, C.B., nlin.CD/0603397
- Diffusively CMLs do have interesting applications in stochastically quantized field theories, in particular for theories that require a **cutoff**.  
C.B., Spatio-temporal Chaos and Vacuum Fluctuations of Quantized Fields, World Scientific (2002)
- A **chaotic dark energy model**, leading to finite vacuum energy, is based on a CML of diffusively coupled 3rd-order Tchebyscheff maps  
C.B., Phys. Rev. D 69, 123515 (2004)