

Non-commutative Field Theory with Twistor-like Coordinates

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Problem

Quantum field theories are singular at short distances ? presence of ultra-violet divergences

These are handled by renormalization (whenever it is possible), often leaving some unpleasant “naturalness” problems

Non-renormalizable theories like quantum gravity are even worse: infinite number of input parameters (UV counterterms) ? no predictive power

So it is a good thing to construct UV softer, or even finite theories - c.f. superstrings, N=8 SUGRA (???) etc.

One way to change the short-distance behavior is to change the “particle” content. SUSY does it by pairing scalars with fermions, superstring theory by upgrading point-like particles to extended objects like strings. What happens in N=8 SUGRA is still unclear...

A more radical and profound idea is to change spacetime, replacing space-time continuum by some discrete or fuzzy “medium”

Very important (and in some way generic) examples of fuzzy spacetimes are those with non-commuting position coordinates

Questions

- *Does spacetime non-commutativity improve short-distance behavior of QFT?*
- *Twistor spacetime is fuzzy – can one think of fuzzy twistors in terms of some non-commutative geometry?*
- *If yes, how does it affect QFT at short (and long) distances?*

Outline

- I. Non-commuting coordinates*
- II. Non-commutative Field Theory on Moyal Plane*
- III. Twistor Theory Revisited*
- IV. Quantum Fields with Twistor-like Coordinates*

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Non-commuting Coordinates

Simplest and tractable example is the Groenewold-Moyal plane R with

$$[x^1; x^0] = i\epsilon^{10}$$

*A radical step – the algebra of functions (fields) on R is modified
– the product is deformed to a star (Moyal) product:*

$$(\hat{A}_1 \star \hat{A}_2)(x) = e^{\frac{i}{2}\epsilon^{10} \partial_y^1 \partial_z^0} \hat{A}_1(y) \hat{A}_2(z) \Big|_{y=z=x}$$

Interesting mathematics. Appears in open string theory in the presence of a constant B-field $B=T$ (Seiberg-Witten, '99). But is it physically sensible?

Non-commutative Field Theory

Is it physically sensible?

Non-commutative Lagrangians involve non-local interactions with star products

$$\int d^4x \left[\frac{1}{2} (\partial_\mu \hat{A})^2 + \frac{1}{2} m^2 \hat{A}^2 + \frac{1}{4!} \hat{A}^4 \right]$$

$$\int d^4x \left[\frac{1}{2} (\partial_\mu \hat{A})^2 + \frac{1}{2} m^2 \hat{A}^2 + \frac{1}{4!} \hat{A} \star \hat{A} \star \hat{A} \star \hat{A} \right]$$

Free Feynman propagators are not affected, but the perturbative interaction vertices are modified by the factors

$$e^{i \sum_{i < j} C_{ij} k_i \wedge k_j} \epsilon^{1\dots}$$

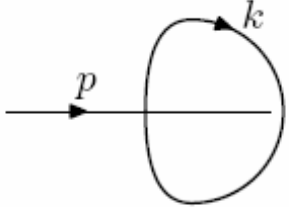
(Minwalla et al, '99)

They affect UV and IR behavior of Feynman diagrams



Non-commutative Field Theory

Is it physically sensible and useful?



$$\begin{aligned}
 &= \frac{\lambda}{12} \int \frac{d^4 k}{(2\pi)^4} \frac{e^{i p_\mu k_\nu \Theta^{\mu\nu}}}{k^2 + m^2} \\
 &= \frac{\lambda}{48\pi^2} \sqrt{\frac{m^2}{(\Theta p)^2}} K_1(\sqrt{m^2(\Theta p)^2}) \underset{p \rightarrow 0}{\simeq} p^{-2}.
 \end{aligned}$$

from Rivasseau, '07

UV-IR mixing: ? 8 , p 0 limits' order matters

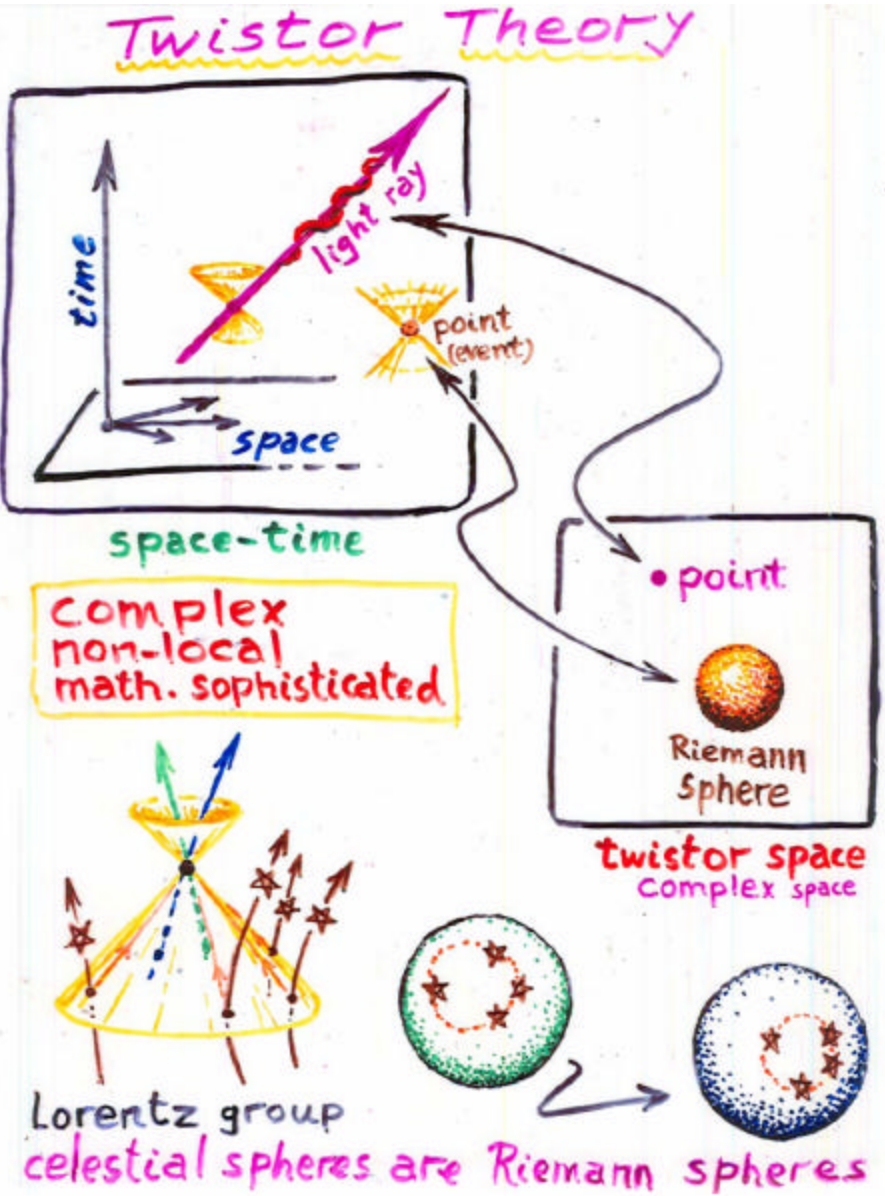
Renormalizable F^2 , without UV/IR mixing, can be constructed by modifying quadratic terms

(Grosse, Wulkenhaar '04)

In general, no significant improvement in UV – Feynman diagrams still have the same degree of divergences as commutative QFT

Moyal Plane non-commutativity isn't too useful for improving short-distance behavior

Twistor Theory (Penrose '67)



Twistors: Spinors representing null geodesics (light rays, world lines) in M ?

Intersections Points

Notation (Penrose, Rindler '86)

$$V^{AA^0} = \begin{pmatrix} V^{00^0} & V^{01^0} \\ V^{10^0} & V^{11^0} \end{pmatrix}$$

$$= \frac{1}{2} \begin{pmatrix} V^0 + V^3 & V^1 + iV^2 \\ V^1 - iV^2 & V^0 - V^3 \end{pmatrix}$$

momentum $p_{AA^0} = \frac{1}{4} A^1 A^0$

$$L^{AA^0 BB^0} = i! (A^1 A^0 B^1 B^0) - i! (A^0 A^1 B^2 B^0)$$

? angular momentum

Twistor Theory

Twistors $Z^{\otimes} = (! A ; \frac{1}{4}A^0)$ $\otimes = 1; 2; 3; 4$

$! A = !^{\pm} A ; i X^{AA^0} \frac{1}{4}A^0 ; \frac{1}{4}A^0 = \frac{1}{4}A^0$
 ? reference point (? line) of angular momentum L

Dual Twistors $\dot{Z}^{\otimes} = (\frac{1}{4}A ; ! A^0)$

$$p^0 = \frac{1}{2}(Z^3 \dot{Z}_1 + Z^2 \dot{Z}_0); \dots$$

$$L^{01} = i ; L^{10} = \frac{i}{2}(Z^0 \dot{Z}_0 ; Z^1 \dot{Z}_1 ; Z^2 \dot{Z}_2 + Z^3 \dot{Z}_3); \dots$$

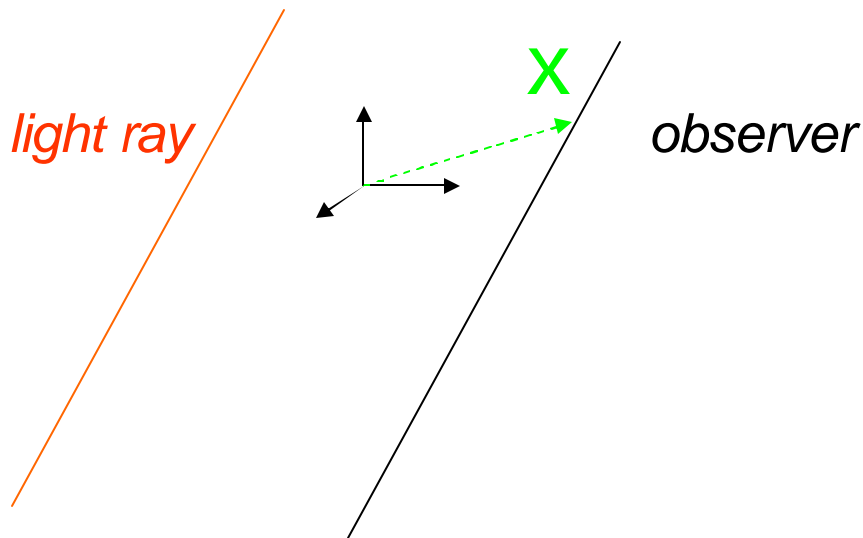
Interpretation

“Angular” twistor $\dot{A} = \dot{A} + iX^{AA^0} \frac{1}{4} A^0$ “incidence relation”

Usually, one considers a fixed reference point, often setting $\dot{A} = 0$
 Then one thinks about x as running along the light ray

$$X^{AA^0} \dot{A} = X^{AA^0} + k \frac{1}{4} A^0$$

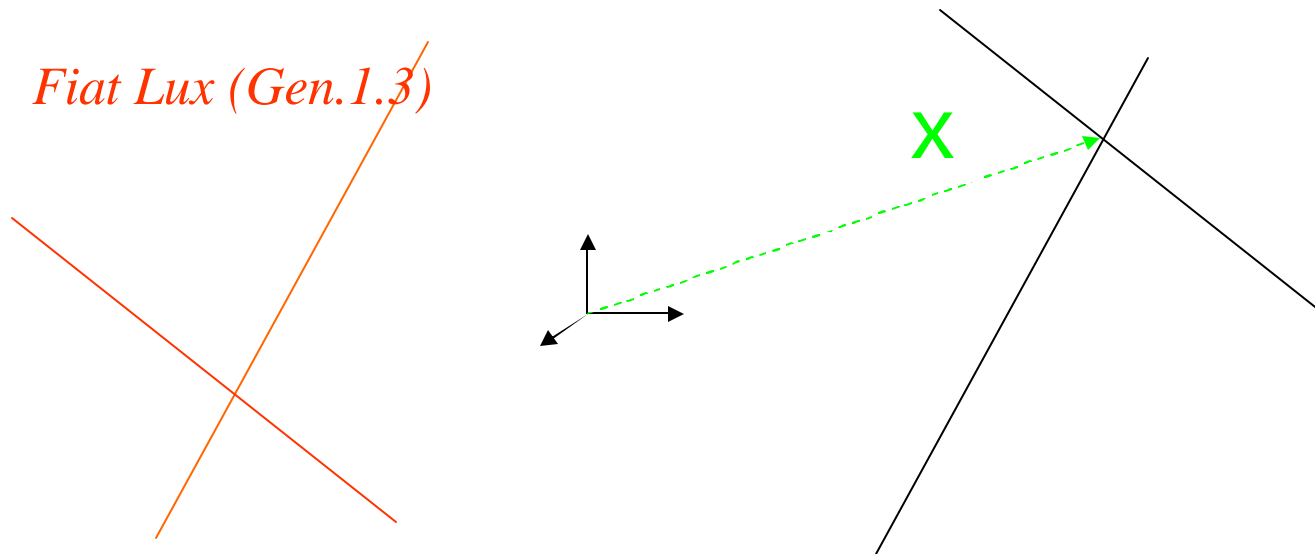
Another viewpoint: consider a fixed light ray :



By using 2 reference light rays, and measuring

$$Z_a^{\textcircled{R}} = \left(\begin{matrix} A \\ a \end{matrix} ; \frac{1}{4} a A^0 \right) \quad a = 1; 2$$

the observer can determine his/her position:



$$X^{AA^0} = \frac{i}{\frac{1}{4} {}_1 B^0 \frac{1}{4} {}_2 B^0} \left[\left(\begin{matrix} A \\ 1 \end{matrix} ; \begin{matrix} A \\ 1 \end{matrix} \right) \frac{1}{4} {}_2 A^0 ; \left(\begin{matrix} A \\ 2 \end{matrix} ; \begin{matrix} A \\ 2 \end{matrix} \right) \frac{1}{4} {}_1 A^0 \right]$$

Twistor Quantization



$$[L_{10}; P_{1/2}] = \sim_{1/2} P_0 \quad ; \quad \sim_{0/2} P_1$$

↓
canonical quantization
of Poincare algebra

$$[Z^{\otimes}; \hat{Z}^-] = \sim_{\pm}^{\otimes}$$

$$[!^A; \frac{1}{4}B] = \sim_{\pm}^A_B$$

$$[\frac{1}{4}A^0; !^B] = \sim_{\pm}^B_{A^0}$$

n.b. units: $! = x^{\rho} \bar{p}$; $\frac{1}{4} = p \bar{p}$

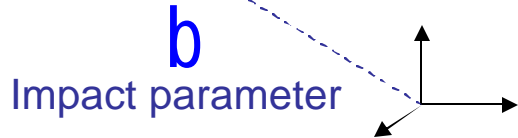
A commonly expressed view:
Quantized metric \rightsquigarrow "fuzzy light cone"



Fuzzy space-time needs no quantum gravity?
What are the consequences
of such non-commutativity?

More on (obvious) interpretation

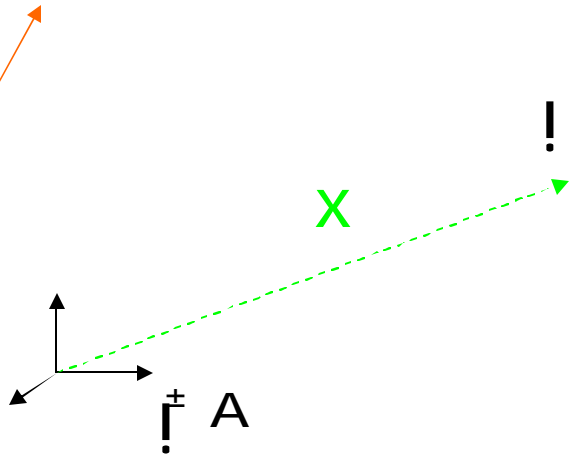
p



$$[!^A; \frac{1}{4}B] = \sim \pm \frac{A}{B} \quad [1/4A^0; !^B] = \sim \pm \frac{B^0}{A^0}$$

$$\Phi X? \Phi p? \sim \frac{\sim}{2}$$

What about angular momenta w.r.t. different points?



$$!^A = i^A; iX^{AA^0} \frac{1}{4}A^0$$

$$\Phi i^A; \Phi \frac{1}{4}A^0; \Phi !$$

$$[!^A(x_1); !^A(x_2)] = i \sim (x_2^{AA^0} i x_1^{AA^0})$$

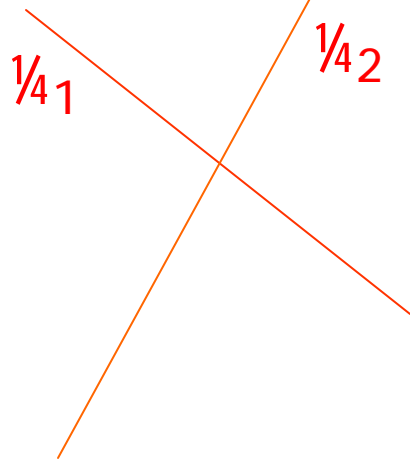
Spacetime parameterized by non-commuting twistor-like position coordinates

$$[! \frac{A}{a}(x_1); ! \frac{A^0}{b}(x_2)] = i \sim (x_2^{AA^0} i x_1^{AA^0})_{\pm ab}$$

LOCALLY COMMUTING BUT

NON-LOCALLY NON-COMMUTING
UNCERTAINTY GROWS WITH SEPARATION (LIKE IN A FOG...)

Reference light rays



LORENTZ SYMMETRY BROKEN BY TIME-LIKE

$$I_{AA^0} = \frac{1}{4}_1 A \frac{1}{4}_1 A^0 + \frac{1}{4}_2 A \frac{1}{4}_2 A^0$$

SO(1,3) ? SO(3)

$$I = (\boldsymbol{\mu} ; \theta)$$

SO IT'S A...

NON-COMMUTATIVITY SCALE

FOGGY ÆTHER

©



Free Fields propagating in *FOGGY ÆTHER*

$$\hat{A}(\mathbf{x}; t) = \int \frac{d^3 p}{(2\pi)^3 2E_p} \left(a_p e^{i(\mathbf{p}\cdot\mathbf{x} - E_p t)} + a_p^\dagger e^{i(\mathbf{p}\cdot\mathbf{x} + E_p t)} \right)$$

SECOND QUANTIZATION:

$$\rho_{AA^0} = \int \frac{d^3 p}{(2\pi)^3} a_p^\dagger a_p \quad [a_p; a_{p^0}^\dagger] = \pm \delta^3(\mathbf{p} - \mathbf{p}^0)$$

FEYNMAN PROPAGATOR:

$$iD(x^0; \mathbf{x}) = \langle 0 | T(\hat{A}(x^0; \mathbf{x}) \hat{A}(0; \mathbf{0})) | 0 \rangle$$



NON-COMMUTATIVE (BAKER-HAUSDORFF)

$$e^A e^B = e^{(A+B + \frac{1}{2}[A;B])}$$

$$e^{i(\mathbf{p}\cdot\mathbf{x} - E_p t)} e^{i(\mathbf{p}^0\cdot\mathbf{x}^0 + E_{p^0} t^0)} = e^{i(\mathbf{p}\cdot\mathbf{x} - E_p t + \mathbf{p}^0\cdot\mathbf{x}^0 + E_{p^0} t^0)}$$

$$= e^{i(\mathbf{p}\cdot\mathbf{x} - E_p t + \mathbf{p}^0\cdot\mathbf{x}^0 + E_{p^0} t^0 + \frac{i}{2}(\mathbf{x} - \mathbf{x}^0)\cdot(\mathbf{p} - \mathbf{p}^0))}$$

Feynman Propagator

$$iD(x^0_i, x) = \int_{\mathbb{R}} \frac{d^3 p}{(2\pi)^3 2|\mathbf{p}|} [e^{i p \cdot (x^0_i - x)} (p^2 + 2i) + e^{i p \cdot (x^0_i - x)} (p^2 - 2i)] \mu(t^0_i, t)$$

$$= i \int_{\mathbb{R}} \frac{d^4 k}{(2\pi)^4} e^{i k \cdot (x^0_i - x)} \mathcal{D}(k)$$

$$\mathcal{D}(k) = \frac{1}{k^2} \left[\frac{1}{|\mathbf{k}|^2 + 1} + \frac{1}{|\mathbf{k}|^2 + 1 + 1} \right]$$

$$\text{UV : } |\mathbf{k}| \gg \mu \quad \mathcal{D}(k) \gg \frac{1}{k^3} \quad \Bigg| \quad \text{IR : } |\mathbf{k}| \ll \mu \quad \mathcal{D}(k) \gg \frac{1}{k^2}$$

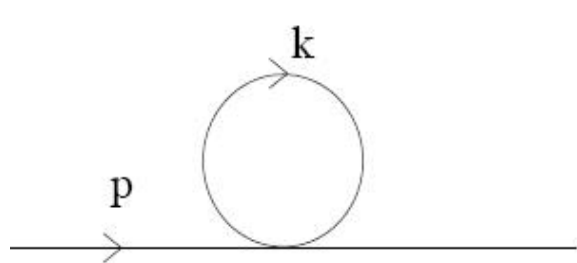
ABOVE μ NON-COMMUTATIVITY SCALE

STANDARD PROPAGATOR

Interacting Fields (very preliminary)

Coordinates are locally commuting ? local interactions unchanged

UV behavior of Feynman diagrams determined by propagators



$$\gg \int \frac{d^4k}{k^3}$$

*Linear (instead of quadratic)
UV divergence*

No UV/IR mixing

gauge theories in FOGGY ÆTHER : perturbatively finite ?

Conclusions

- Non-local (foggy) non-commutativity in twistor space
- Determined by 2 constants: h, μ
- Lorentz symmetry broken: preferred time direction, fundamental time unit $\zeta = \sim (1/c^2)$ and rotational invariance in the \mathcal{A} ether frame
- Assuming \mathcal{A} ether = CMB ? $\mu > 10 \text{ TeV}$
- Foggy spacetime tames UV divergences of QFT, no UV/IR mixing
- Many open questions: interacting QFT formalism, divergences, gravity, ...



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