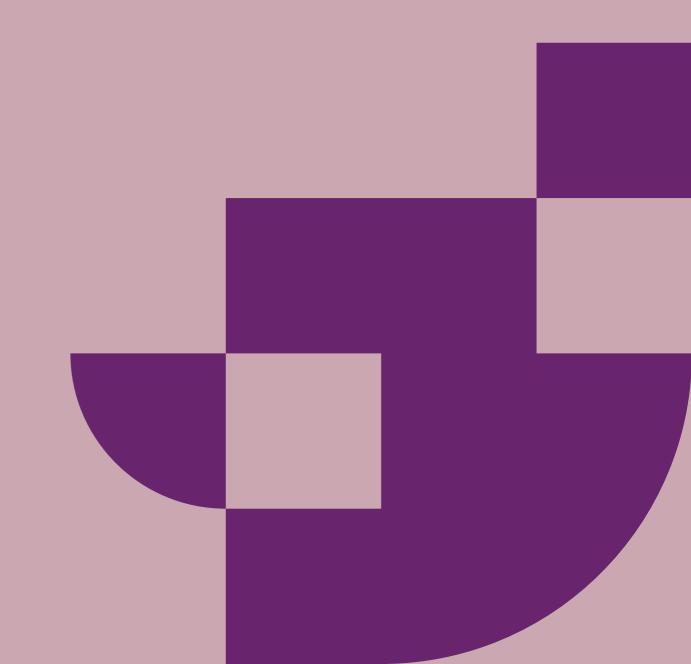


Revisiting Taylor relaxation

Anthony Yeates

with Alexander Russell (Dundee)

BAMC, Glasgow, 7-Apr-2021



Aims

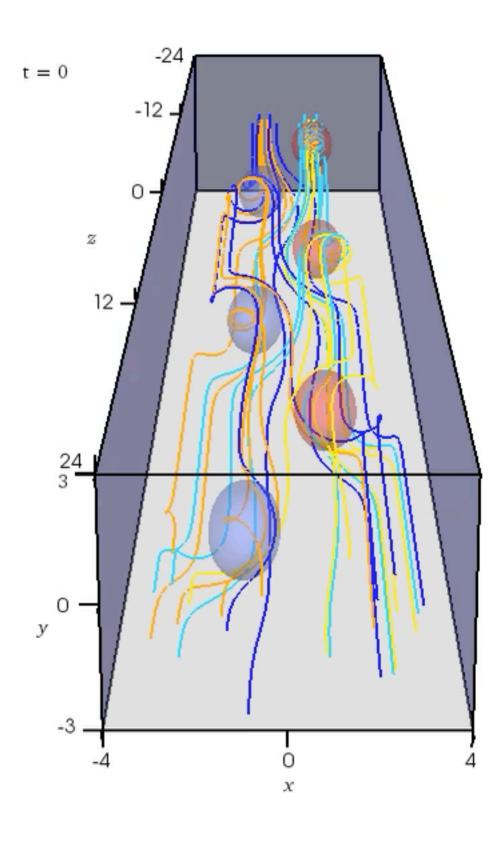
- Resistive relaxation of plasma with initially "braided" magnetic field shows spontaneous self-organization.
- ▶ Relaxed state not predicted by **Taylor relaxation** theory: conservation of total magnetic helicity => linear force-free field $\nabla \times \mathbf{B} = \lambda_0 \mathbf{B}$.

[Taylor, Rev Mod Phys 58, 741, 1986]

Can we learn more by following the evolution of field line helicities?

$$\mathcal{A}(L) = \lim_{\epsilon o 0} rac{\int_{V_{\epsilon}(L)} \mathbf{A} \cdot \mathbf{B} \, \mathrm{d}V}{\Phi(V_{\epsilon}(L))} = \int_{L} \mathbf{A} \cdot \, \mathrm{d}\mathbf{I}$$

[Berger, *Astron Astrophys* **201**, 355, 1988; Yeates & Hornig, *Phys Plasmas* **20**, 012102, 2013; Aly, *Fluid Dyn Res* **50**, 011408, 2018]



▶ Taylor assumed that the **A.B** would be arbitrarily redistributed between field lines so individual field line helicities play no role in determining the final state.

Simulation setup

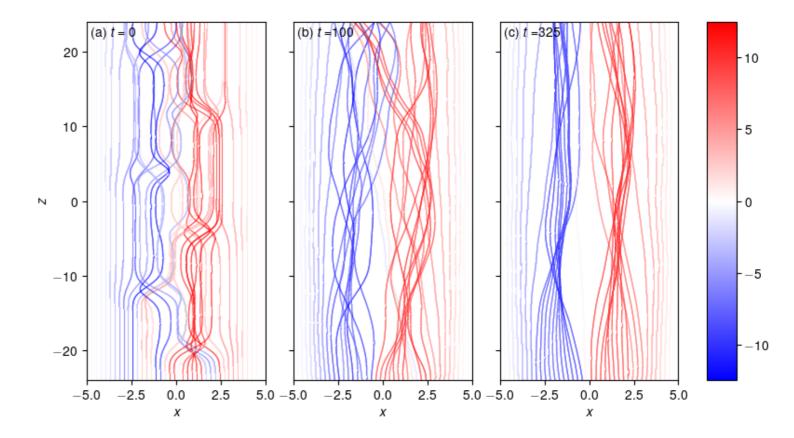
▶ Resistive-MHD equations in Cartesian domain [-8,8] x [-8,8] x [-24, 24].

$$\begin{split} &\frac{\partial \rho}{\partial t} = -\nabla \cdot (\rho \mathbf{v}) \\ &\rho \frac{\mathsf{D} \mathbf{v}}{\mathsf{D} t} = \mathbf{j} \times \mathbf{B} - \nabla \rho + [\mathsf{viscosity}] \\ &\rho \frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B}) - \nabla \times (\eta \mathbf{j}) \\ &\rho \frac{\mathsf{D} \epsilon}{\mathsf{D} t} = -\rho \nabla \cdot \mathbf{v} + \eta |\mathbf{j}|^2 + [\mathsf{viscous \ disspn.}] \\ &\rho = \rho \epsilon (\gamma - 1) \qquad \qquad \beta \approx 0.01 \qquad \gamma = \frac{5}{3} \\ &\mu_0 \mathbf{j} = \nabla \times \mathbf{B} \end{split}$$

- Lundquist number $S = \eta^{-1}$ from 2,500 to 20,000.
- ▶ Line-tied boundaries v = 0
- Initially braided magnetic field.
- ▶ LARE3d code (T. Arber). https://github.com/Warwick-Plasma/Lare3d

Evolution of field line helicity

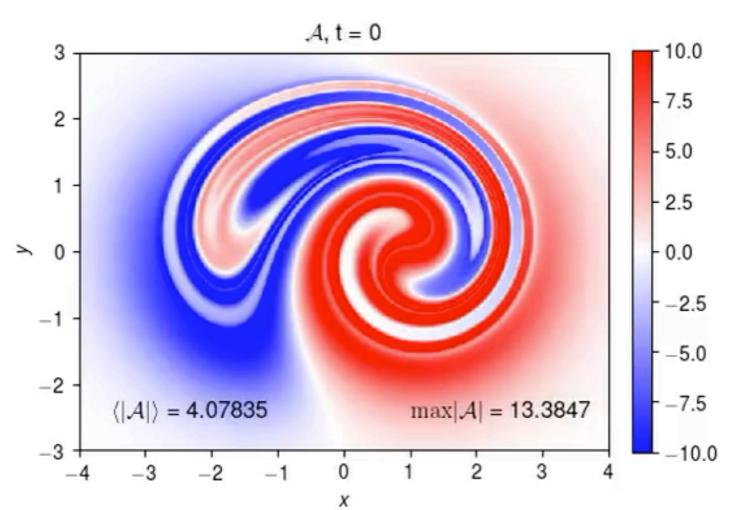
► Compute A and integrate along field lines for a sequence of snapshots.

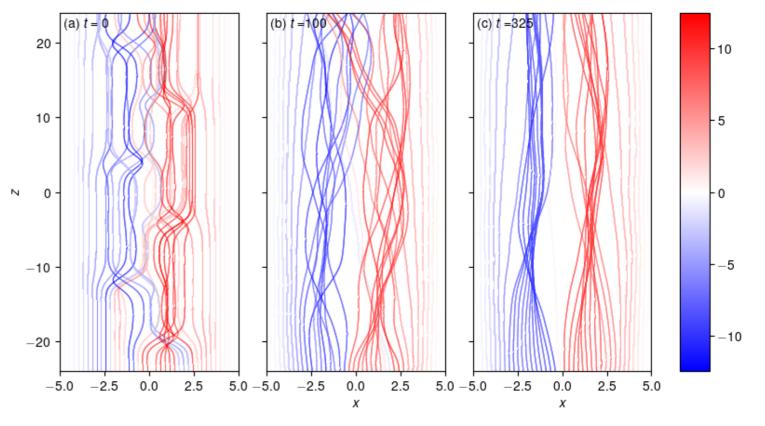


Evolution of field line helicity

▶ Compute A and integrate along field lines for a sequence of snapshots.

Cross-section on the lower boundary:





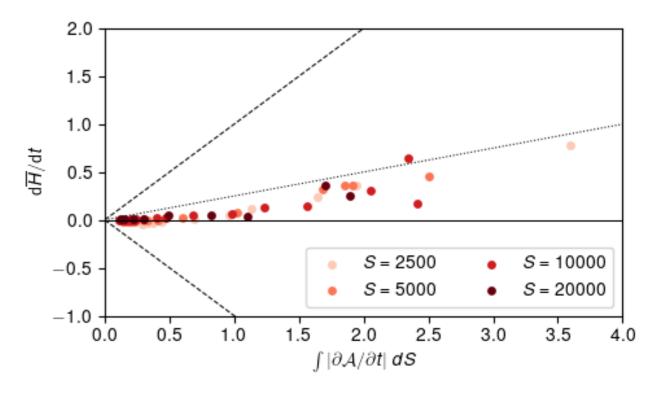
- 1. To leading order, field line helicity is redistributed rather than destroyed.
- 2. The relaxed state exhibits **self organization** into distinct positive and negative regions.
- 3. Within each of these regions, the field line helicity is strikingly **uniform**.

1. Dominance of redistribution

Verifies our earlier prediction that evolution of field line helicity in high Rm is dominated by redistribution rather than dissipation.

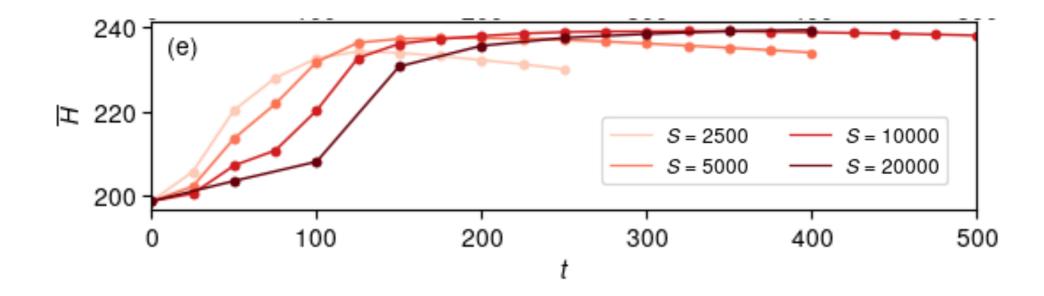
[Russell et al., Phys Plasmas 22, 032106, 2015]

$$\overline{H} = \int_{-4}^{4} dx \int_{-4}^{4} dy \left(|\mathcal{A}| B_z \right)_{z=-24}$$



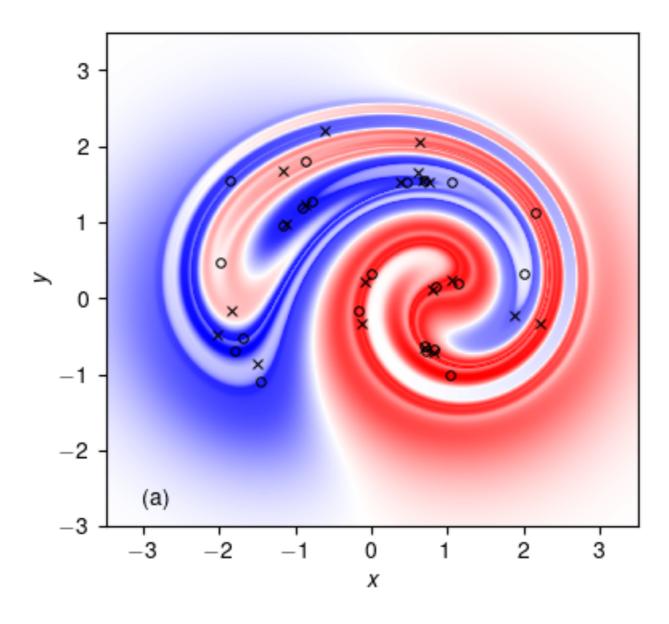
[See upcoming paper for details of the evolution equation that led to this prediction.]

Notice that the small changes in \overline{H} do have a net positive drift:



2. Self organization

▶ Changes in FLH are localized, so the total Poincaré index of critical points (where $\nabla A = \mathbf{0}$) is invariant.



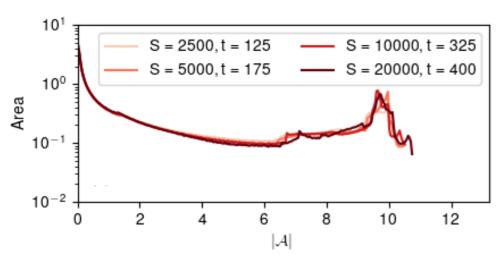
▶ Initially +2 (+22, -20) — explains two final regions (one maximum, one minimum).

3. Uniformization

Field line helicity is much more uniform than the "force-free parameter",

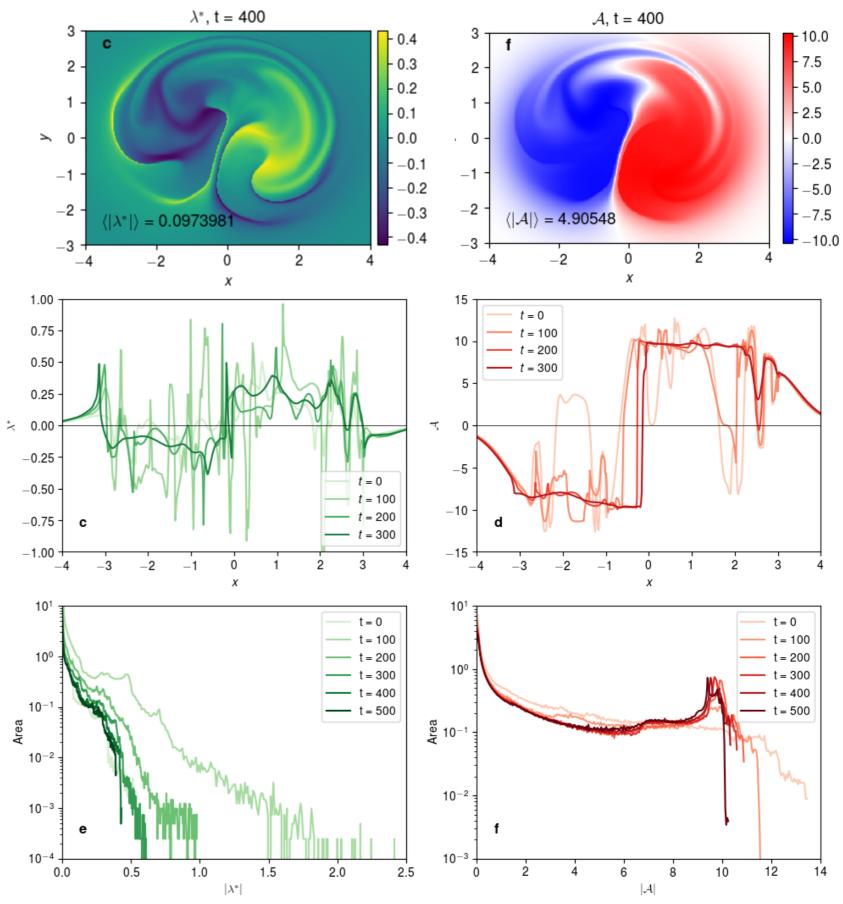
$$\lambda = \frac{\mathbf{j} \cdot \mathbf{B}}{|\mathbf{B}|^2}$$

The relaxed state appears independent of Lundquist number.





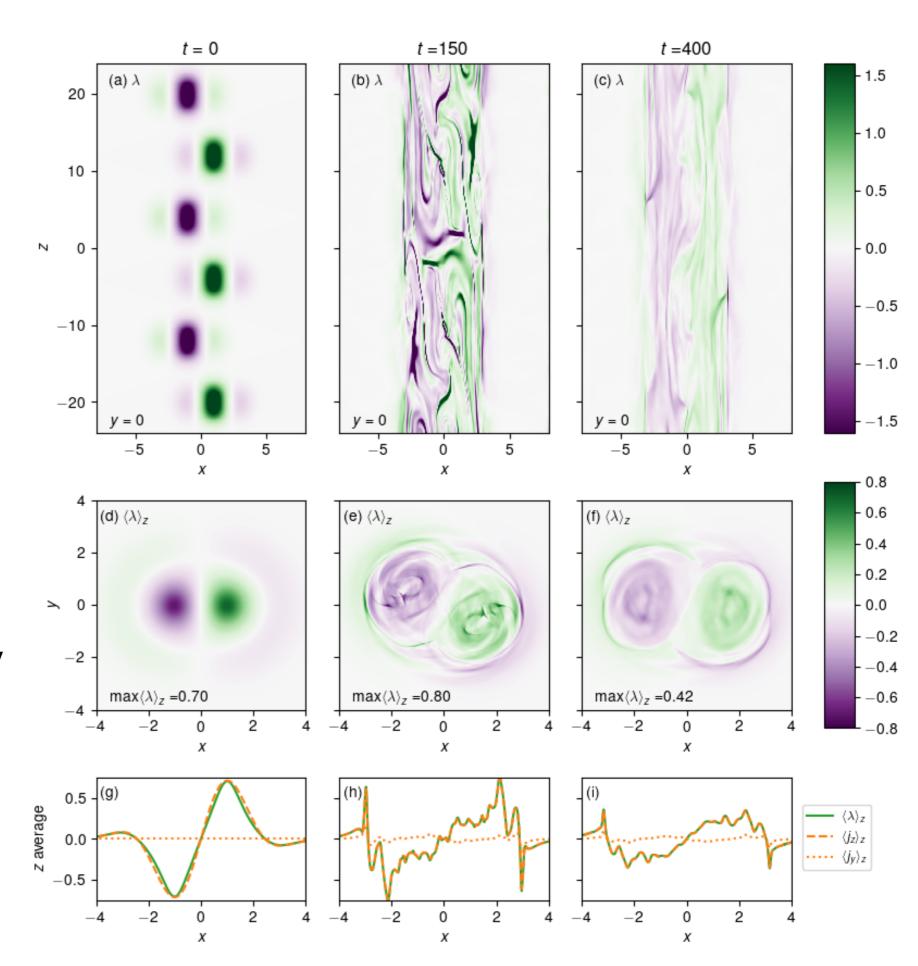
Field line helicity



Hypothesis

- There is a tendency for Taylor relaxation within each tube, with λ ≈ constant to (rough) first approximation.
- With our strong guide field, $\lambda \approx j_z$.
- An axisymmetric field with constant j_z has uniform field line helicity

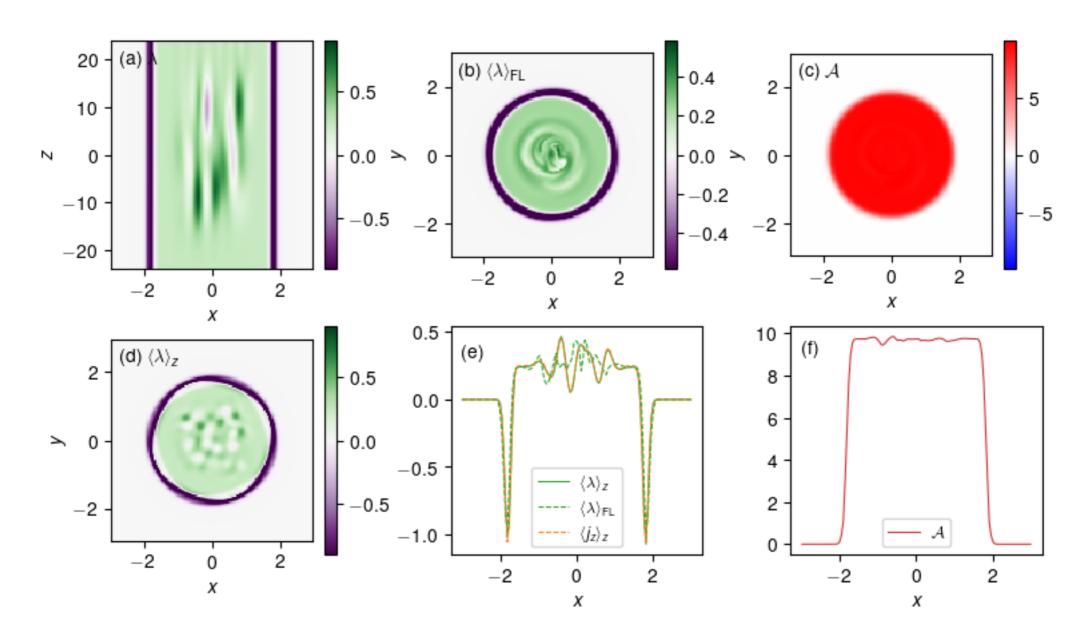
$$\mathcal{A}=\frac{j_zR^2L}{4}.$$



In the presence of local fluctuations on a uniform twist tube, field line helicity is more robust than λ because it is a **global average quantity**.

[cf. Prior & Yeates, Astrophys J 787, 100, 2014]

▶ Simple model — uniform twist + fluctuations:



Conclusions

- ▶ Field line helicities are not arbitrarily changed but are **redistributed** more effectively than increased/decreased.
- ▶ **Self-organization** into a relaxed state with two discrete magnetic flux tubes may be predicted from the initial field line helicity distribution (but not from the total helicity).
- Within each tube, the final state is best described as a state of uniform field line helicity, independent of Lundquist number. This arises even from an "approximate" Taylor relaxation where the force-free parameter is rather less uniform.

A.J.B. Russell, A.R. Yeates, G. Hornig & A.L. Wilmot-Smith, Evolution of field line helicity during magnetic reconnection, *Phys Plasmas* **22**, 032106, 2015.

A.R. Yeates & A.J.B. Russell, Evolution of field line helicity in magnetic relaxation, in preparation.

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